## Chapter 3

## RIEMANNIAN MANIFOLDS

In his ground-breaking essay in 1855, Riemann investigated the nature of a space $M$ (called $n$ extended quantities by him) through a precise mechanism of measurement of length. Among others, Riemann studied in details the fundamental case when the measurement takes the form

$$
d s=\sqrt{\sum g_{i j} d x^{i} d x^{j}},
$$

now called a Riemannian metric $g$. Through local expansion, Riemann discovered the invariant quantity $R_{i j k l}$ in the second order terms and proved that its vanishing is equivalent to that $(M, g)$ is locally Euclidean.

The Riemann curvature tensor $\sum R_{i j k l} d x^{i} \otimes d x^{j} \otimes d x^{k} \otimes d x^{l}$, which generalizes the Gaussian curvature $K$ of surfaces, was (and still is) difficult to work with. To understand it, tensor calculus was subsequently developed in the next decades by Levi-Civita, Christoffel, Ricci and others. Notably, for each $X \in T_{p} M$ the notion of absolute differentiation $\nabla_{X} T$ was invented. Aided by this language, Hilbert and Einstein wrote down the field equation of General Relativity in 1915, which confirmed the central role of Riemannian geometry in Geometry and Physics till today.

In this chapter we present the basic elements in Riemannian geometry. We introduce normal coordinates (local) and exponential maps (global) and prove the Hopf-Rinow theorem to characterize complete manifolds. We then discuss geodesics and variations of them. The precise link between the second variation formula and the curvature tensor leads to topological applications like Bonnet and Synge's theorems. Together with Jacobi fields we prove the Cartan-Hadamard theorem which characterizes spaces with constant curvature as quotients of $\mathbb{R}^{N}, S^{N}$ or $\mathbb{H}^{N}$ (space forms).

We discuss also the second fundamental form $B$ and the mean curvature vector $H$ of a submanifold $i: M \hookrightarrow N$ and study the variational aspects. Comparing with geodesics, only the simplest properties are addressed here. More advanced aspects will be discussed in later chapters.

## 1. Riemannian structures and affine structures

Definition 3.1. A Riemannian manifold $(M, g)$ is a $C^{\infty}$ manifold $M$ together with $g \in C^{\infty}\left(M, \operatorname{Sym}^{2}\left(T^{*} M\right)\right)$ such that $g_{p}$ is positive definite for every $p \in M$.

Such a $g$ is called a Riemannian metric, or simply a metric. In local chart $(U, \mathbf{x})$,

$$
g=\sum_{i, j} g_{i j} d x^{i} \otimes d x^{j}
$$

where $g_{i j}=g\left(\frac{\partial}{\partial x^{i}}, \frac{\partial}{\partial x^{j}}\right)$ on $U$. We also denote $\langle X, Y\rangle=g(X, Y)$.
Any differentiable manifold $M$ can be endowed with Riemannian metrics through a P.O.U. $\left\{U_{\alpha}, \rho_{\alpha}\right\}$ : we can assign an arbitrary metric $g_{\alpha}$ on $U_{\alpha}$ (e.g. the Euclidean inner product), and then define

$$
g=\sum_{\alpha} \rho_{\alpha} g_{\alpha}
$$

It is easily to see that $g$ is a metric. It is also easy to give another construction of metrics on $M$ under a given $C^{\infty}$ imbedding $M \hookrightarrow$ $\mathbb{R}^{N}$. We simply consider the inner product on $T_{p} M \hookrightarrow T_{p} \mathbb{R}^{N} \cong \mathbb{R}^{N}$ induced from the one on $T_{p} \mathbb{R}^{N}$, say the Euclidean one $g_{\mathbb{R}^{N}} .{ }^{1}$

With a Riemannian metric $g$, we can define the concept of measure and integration even for non-orientable $M$ : for $f \in C_{c}^{\infty}(M)$,

$$
\int_{M} f:=\int_{M} f d \mu_{g}
$$

where the measure, or the volume form,

$$
d \mu_{g} \equiv d V \equiv d \mathrm{vol}:=|\omega|
$$

[^0]is the absolute value of a top form $\omega$ defined at any $p \in M$ by
$$
\omega_{p}:=\phi^{1} \wedge \ldots \wedge \phi^{n}
$$
where $\phi^{n}, \ldots, \phi^{n}$ of $T_{p}^{*} M$ is any orthonormal basis of $T_{p}^{*} M$, say dual to an orthonormal basis $e_{1}, e_{2}, \ldots, e_{n}$ of $T_{p} M$.

Exercise 3.1. (2) Given $v_{1}, \ldots, v_{k} \in V \cong \mathbb{R}^{n}$ with inner product $\langle$,$\rangle ,$ show that the $k$-dimensional volume of the parallelpipade

$$
P\left(v_{1}, \ldots, v_{k}\right)=\left\{\sum_{i=1}^{k} t_{i} v_{i} \mid 0 \leq t_{i} \leq 1\right\}
$$

is $\sqrt{\operatorname{det}\left(\left\langle v_{i} \cdot v_{j}\right\rangle\right)_{i, j=1}^{k}}$.
(2) On a local chart $(U, \mathbf{x})$, show that

$$
d \mu_{g} \equiv \omega=\sqrt{\operatorname{det}\left(g_{i j}\right)} d x^{1} \wedge \ldots \wedge d x^{n}
$$

Remark 3.2. In order for the volume form to be defined we need only the manifold and the metric $g$ to be $C^{1}$.

With a metric $g$, we can also measure the size of lower dimensional submanifolds. Given a $C^{1}$ immersion $\iota: S \rightarrow M$ of a $k \leq n$ dimensional manifold $S$, the induced tensor $g_{S}:=\iota^{*} g$ is a Riemannian metric on $S$ and then the $k$ dimensional volume form $d \mu_{g_{S}} \in A^{k}(S)$ is defined. In local coordinates $\left(x^{1}, \ldots, x^{n}\right)$ on $M$ and $\left(t^{1}, \ldots, t^{k}\right)$ on $S$, we see from Exercise 3.1 that

$$
\begin{aligned}
d \mu_{g_{S}} & =\sqrt{\operatorname{det}\left(g\left(\iota_{*} \partial_{t^{i}, \iota_{*}} \partial_{t j}\right)\right)_{i, j=1}^{k}} d t^{1} \wedge \ldots \wedge d t^{k}, \\
& =\sqrt{\operatorname{det}\left(\sum_{i^{\prime}, j^{\prime}=1}^{n} g_{i^{\prime} j^{\prime}} \frac{\partial x^{i^{\prime}}}{\partial t^{i}} \frac{\partial x^{j^{\prime}}}{\partial t^{j}}\right)_{i, j=1}^{k}} d t^{1} \wedge \ldots \wedge d t^{k}, \\
& =\sqrt{\operatorname{det}(D \tilde{\iota})^{t} G(D \tilde{\iota})} d t^{1} \wedge \ldots \wedge d t^{k},
\end{aligned}
$$

where $\partial_{t^{i}}:=\frac{\partial}{\partial t^{i}}, \iota_{*} \partial_{t^{i}}=\sum_{i^{\prime}} \frac{\partial x^{i^{\prime}}}{\partial t^{i}} \frac{\partial}{\partial x^{i^{\prime}}}$, and $G=\left(g_{i^{\prime} j^{\prime}}\right)$. The case $k=1$ goes back to arc length element we start with:

$$
d s^{2}=\sum g_{i j} d x^{i} d x^{j}
$$

Next, we study differentiations on $(M, g)$. Let $F$ be a vectorvalued function on $\mathbb{R}^{n}$ and $v$ a unit vector. The directional derivative of $F$ at a point $p$ in the $v$ direction is defined by

$$
D_{v} F:=\lim _{t \rightarrow 0} \frac{1}{t}(F(p+t v)-F(p)) .
$$

To generalize it to a manifold $M$, we need to compare the difference of vectors (or tensors) at different points. Therefore, we need to "parallel translate" them to the same point $p$ to compare the difference. As a result, to define the differentiation, what we really need is the notion of parallel translations or affine structures.

We wish to define the directional derivative such that it only depends on $v \in T_{p} M$ but not on the extension $\tilde{v}$ of it. Also, the resulting differentiation should satisfy linearity and the Leibniz rule.

The "Lie" derivative does not meet the requirement. From

$$
L_{f X} Y=[f X, Y]=f[X, Y]-(Y f) X,
$$

we see that it is not function-linear in $X$ and the differentiation of $Y$ at $p$ in the direction $X_{p}$ depends on the extension of $X_{p}$ to $X .{ }^{2}$

The correct concept, developed by Christoffel, Levi-Civita, Ricci, Koszul and others, start from the following simple definition:

Definition 3.3. A covariant differentiation is an operator

$$
\nabla_{X}: C^{\infty}(T M) \rightarrow T_{p} M
$$

for every $p \in M$ and $X \in T_{p} M$, such that
(1) $\nabla_{X} Y$ is $\mathbb{R}$-linear in both $X$ and $Y$, and
(2) (Leibniz rule) $\nabla_{X}(f Y)=X(f) Y_{p}+f(p) \nabla_{X} Y$ for $f \in C_{p}^{\infty}$.

In other words, by the Fundamental Theorem of Tensor Calculus (theorem 2.14), the $\mathbb{R}$ bi-linear operator

$$
\begin{aligned}
\nabla: \quad C^{\infty}(T M) \times C^{\infty}(T M) & \longrightarrow C^{\infty}(T M) \\
(X, Y) \longmapsto & \nabla_{X} Y
\end{aligned}
$$

[^1]
is function-linear in $X$. Thus $\nabla Y$ is a vector valued 1-form:
$$
\nabla: C^{\infty}(T M) \rightarrow A^{1}(T M)=A^{1}(M) \otimes C^{\infty}(T M)
$$

More generally, for a vector bundle $E \rightarrow M$, an affine connection is an operator $D: C^{\infty}(E) \rightarrow A^{1}(E)$ with Leibniz rule. Thus a covariant differentiation is simply an affine connection on $T M$.

Example 3.4. (1) $M=\mathbb{R}^{n}, \nabla_{X}=D_{X}$, the directional derivative.
(2) Let $M \hookrightarrow \mathbb{R}^{n}$ be a submanifold and $p \in M$. Let $X \in T_{p} M$ and $Y$ be a local vector field on $M$ defined near $p$. Let $\widetilde{Y}$ be any $C^{\infty}$ extension of $Y$ to $T \mathbb{R}^{n}$. Then we define the induced connection by

$$
\nabla_{X} Y:=\left(D_{X} \widetilde{Y}\right)^{T}
$$

the tangential part of $D_{X} \widetilde{Y}$ in (1) under the orthogonal projection

$$
T_{p} \mathbb{R}^{n}=T_{p} M \stackrel{\perp}{\oplus} N_{p}
$$

On a general manifold $M$, affine connections on a vector bundle $E \rightarrow M$ can be constructed by patching the local connections. Let $M=\bigcup_{\alpha} U_{\alpha}$ with a partition of unity $\left(U_{\alpha}, \phi_{\alpha}\right)^{\prime}$ s. Given an affine connection $\nabla^{\alpha}$ on $U_{\alpha}$ for each $\alpha$, it is natural to ask if

$$
\nabla:=\sum_{\alpha} \phi_{\alpha} \nabla^{\alpha}
$$

is a covariant differentiation on the whole $M$ ?
Notice that a simple scaling of a connection is no longer a connection. For $\widetilde{\nabla}=h \nabla$,

$$
\widetilde{\nabla}_{X}(f Y)=h X(f) Y+h f \nabla_{X} Y=h X(f) Y+f \widetilde{\nabla}_{X} Y
$$

Leibniz rule fails unless $h=1$. Nevertheless, for a (locally) finite sum $\nabla=\sum h_{i} \nabla^{i}$, the same calculation shows that

$$
\nabla_{X}(f Y)=\left(\sum h_{i}\right) X(f) Y+f \nabla_{X} Y
$$

Thus $\nabla$ is a connection if and only if $\sum_{i} h_{i}=1$, i.e. $\nabla$ is a an affine linear combination of $\nabla^{i}$ s. In particular, for a P.O.U. $\left\{\phi_{\alpha}\right\}, \nabla=\sum_{\alpha} \phi_{\alpha} \nabla^{\alpha}$ is a global covariant differentiation operator.

Exercise 3.2. Show that the set of all affine connections on $E \rightarrow M$ is an affine space $\nabla^{0}+A^{1}$ (End $E$ ) modeled on the infinite dimensional vector space $A^{1}($ End $E)$, where $\nabla^{0}$ is any affine connection.

So far, we have not made use of the Riemannian structure $g$. When $g$ is present, which connection(s) will be the best choice(s) among the above infinite-dimensionally many ones?

Notice that in the case $M \hookrightarrow \mathbb{R}^{n}$ as in Example 3.4, two more properties hold for the induced connection $\nabla$ :
(1) $\nabla$ is compatible with metric, i.e. it satisfies the "Leibniz rule":

$$
\begin{aligned}
& X\langle Y, Z\rangle=\left\langle D_{X} Y, Z\right\rangle+\left\langle Y, D_{X} Z\right\rangle=\left\langle\nabla_{X} Y, Z\right\rangle+\left\langle Y, \nabla_{X} Z\right\rangle \\
& \quad \text { since }\left\langle\left(D_{X} Y\right)^{N}, Z\right\rangle=0 \text { for } Z \in C^{\infty}(T M) .
\end{aligned}
$$

(2) $\nabla$ satisfies the torsion-free condition:

$$
\nabla_{X} Y-\nabla_{Y} X=(X Y-Y X)^{T}=[X, Y]^{T}=[X, Y]
$$

Definition 3.5 (Torsion tensor). For any affine connection $\nabla$ on $T M$, $T(X, Y):=\nabla_{X} Y-\nabla_{Y} X-[X, Y]$ is a tensor, called the torsion of $\nabla$.

Exercise 3.3. Let $(N, g)$ be a Riemannian manifold and $M \hookrightarrow N$. Show that the induced connection $\nabla_{X} Y:=\left(D_{X} \widetilde{Y}\right)^{T}$ is (i) independent of the choices of $\widetilde{Y}$, (ii) compatible with metric and torsion-free.

A fundamental result due to Christoffel (1869) and later used by Levi-Civita (1917) to study parallel transport, asserts that there is an unique connection, the Levi-Civita connection, satisfies (1) and (2):

Theorem 3.6 (Fundamental Theorem of Riemannian Geometry).
For any Riemannian manifold $(M, g)$, there is a unique connection $\nabla=\nabla^{L C}$ on TM which is metric compatible and torsion-free.

Proof of uniqueness. Given any 3 vector fields $X, Y, Z$, by compatibility and torsion-free condition, we have

$$
\begin{aligned}
X\langle Y, Z\rangle & =\left\langle\nabla_{X} Y, Z\right\rangle+\left\langle Y, \nabla_{X} Z\right\rangle \\
Y\langle X, Z\rangle & =\left\langle\nabla_{Y} Z, X\right\rangle+\left\langle Z, \nabla_{Y} X\right\rangle \\
& =\left\langle\nabla_{Y} Z, X\right\rangle+\left\langle Z, \nabla_{X} Y+[Y, X]\right\rangle \\
Z\langle X, Y\rangle & =\left\langle\nabla_{Z} X, Y\right\rangle+\left\langle X, \nabla_{Z} Y\right\rangle \\
& =\left\langle\nabla_{X} Z+[Z, X], Y\right\rangle+\left\langle X, \nabla_{Y} Z+[Z, Y]\right\rangle
\end{aligned}
$$

By adding the first two and subtracting the third one, we get

$$
\begin{aligned}
\left\langle\nabla_{X} Y, Z\right\rangle= & \frac{1}{2}(X\langle Y, Z\rangle+Y\langle Z, X\rangle-Z\langle X, Y\rangle \\
& +\langle Z,[X, Y]\rangle+\langle Y,[Z, X]\rangle-\langle X,[Y, Z]\rangle) .
\end{aligned}
$$

Since $Z$ is arbitrary, this implies that $\nabla_{X} Y$ is uniquely determined by the differentiable structure and the metric.

Exercise 3.4. Prove the existence of Levi-Civita connection by checking the above formula for $\nabla_{X} Y$ defines a covariant differentiation.

In local coordinate $\left(x^{1}, \ldots, x^{m}\right)$, let $\partial_{i}:=\frac{\partial}{\partial x^{i}}$. The Christoffel symbol $\Gamma_{i j}^{k}$ of an affine connection $\nabla$ is defined by

$$
\nabla_{\partial_{i}} \partial_{j}=: \sum_{k} \Gamma_{i j}^{k} \partial_{k}
$$

It measures how the frame moves, hence determines the information of the affine connection: for any vector field $v=\sum v^{j} \partial_{j}$,

$$
\begin{aligned}
\nabla_{i} v & \equiv \nabla_{\partial_{i}}\left(\sum_{j} v^{j} \partial_{j}\right)=\sum_{j}\left(\partial_{i} v^{j}\right) \partial_{j}+\sum_{j, k} v^{j} \Gamma_{i j}^{k} \partial_{k} \\
& =\sum_{j}\left(\partial_{i} v^{j}+\sum_{k} v^{k} \Gamma_{i k}^{j}\right) \partial_{j}=: \sum_{j} v_{; i}^{j} \partial_{j} .
\end{aligned}
$$

Here $v_{; i}^{j}=\partial_{i} v^{j}+\sum_{k} v^{k} \Gamma_{i k}^{j}$. We will generalize this to all tensors later. Lemma 3.7. An affine connection $\nabla$ is torsion-free if and only if $\Gamma_{i j}^{k}=\Gamma_{j i}^{k}$ in all coordinate systems.

Indeed, if $\nabla$ is torsion-free, then

$$
0=\left[\partial_{i}, \partial_{j}\right]=\nabla_{\partial_{i}} \partial_{j}-\nabla_{\partial_{j}} \partial_{i}=\left(\Gamma_{i j}^{k}-\Gamma_{j i}^{k}\right) \partial_{k} .
$$

Conversely, the torsion tensor $T(X, Y)=0$ since $T\left(\partial_{i}, \partial_{j}\right)=0$.
We further assume that $\nabla=\nabla^{L C}$ is the Levi-Civita connection. Since $\left\langle\partial_{i}, \partial_{j}\right\rangle=g_{i j}$, the formula in the proof of Theorem 3.6 says

$$
\sum_{\ell} \Gamma_{i j}^{\ell} g_{\ell k}=\frac{1}{2}\left(\partial_{i} g_{j k}+\partial_{j} g_{k i}-\partial_{k} g_{i j}\right)
$$

Denote by $\left(g^{i j}\right)=\left(g_{i j}\right)^{-1}$, i.e. $g^{p q} g_{q r}=\delta_{r}^{p}$. Then we arrive at the fundamental formula for the Christoffel symbols:

Lemma 3.8. The Christoffel symbols for $\nabla^{L C}$ are given by

$$
\begin{equation*}
\Gamma_{i j}^{k}=\frac{1}{2} \sum_{\ell} g^{k \ell}\left(\partial_{i} g_{j \ell}+\partial_{j} g_{\ell i}-\partial_{\ell} g_{i j}\right) \tag{3.1}
\end{equation*}
$$

Now we turn to the question of parallel translation for an affine connection $\nabla$. Given a smooth curve $\gamma:[0,1] \rightarrow M$. Denote $\dot{\gamma}(t)=$ $\gamma^{\prime}(t)=d \gamma\left(\frac{\partial}{\partial t}\right) \equiv \frac{\partial}{\partial t}=\partial_{t}$. Let $F$ be a vector field along $\gamma$. Write

$$
\frac{D F}{d t}=\nabla_{\partial_{t}} F=\nabla_{t} F
$$

as the covariant differentiation along $\gamma$.
Definition 3.9. We say that $F$ is parallel along $\gamma$ if $\nabla_{\partial_{t}} F=0$.


Let $F=\sum f^{i} \partial_{i}, \gamma(t)=\left(x^{1}(t), \ldots, x^{n}(t)\right)$. Then

$$
\begin{aligned}
\nabla_{\partial_{t}} F & =\sum\left(\frac{\partial f^{i}}{\partial t} \partial_{i}+f^{i} \nabla_{\partial_{t}} \partial_{i}\right)=\sum\left(\frac{\partial f^{i}}{\partial t} \partial_{i}+f^{i} \dot{x}^{j} \Gamma_{i j}^{k} \partial_{k}\right) \\
& =\sum\left(\frac{\partial f^{i}}{\partial t}+f^{k} \dot{x}^{j} \Gamma_{k j}^{i}\right) \partial_{i} .
\end{aligned}
$$

The equation $\nabla_{t} F=0$ is a first order linear ODE in $t$. Thus the parallel translation $F(t)$ exists along $\gamma$ and it is uniquely determined by the initial vector $F(0)$.

Lemma 3.10. Let $(M, g)$ be a Riemannian manifold. If $\nabla$ is metric compatible then parallel translation preserves inner products.

Proof. Let $V_{1}, V_{2}$ be two parallel vector fields along $\gamma$. Then

$$
\dot{\gamma}\left\langle V_{1}, V_{2}\right\rangle=\left\langle\nabla_{\dot{\gamma}} V_{1}, V_{2}\right\rangle+\left\langle V_{1}, \nabla_{\dot{\gamma}} V_{2}\right\rangle=0 .
$$

Thus $\left\langle V_{1}, V_{2}\right\rangle$ is a constant function in $t$.
2. Geodesics, exponential map and Riemann's normal coordinates

From now on we assume that $\nabla=\nabla^{L C}$.

## Definition 3.11. A (parametrized) curve $\gamma$ is a geodesic if $\nabla_{\dot{\gamma}} \dot{\gamma}=0$.

Example 3.12. For $M \hookrightarrow \mathbb{R}^{n}$,

$$
\nabla_{\dot{\gamma}} \dot{\gamma}=\left(D_{\dot{\gamma}} \dot{\gamma}\right)^{T}=\left(D_{\partial_{t}} \dot{\gamma}\right)^{T}=\gamma^{\prime \prime}(t)^{T} .
$$

In particular, geodesics on linear spaces are lines and geodesics on spheres $S^{m}$ are great circles.

If $\nabla_{\dot{\gamma}} \dot{\gamma}=0$ then

$$
\dot{\gamma}\langle\dot{\gamma}, \dot{\gamma}\rangle=2\left\langle\nabla_{\dot{\gamma}} \dot{\gamma}, \dot{\gamma}\right\rangle=0 .
$$

So $|\dot{\gamma}|=c$ is a constant and the arc length $s(t)=\int_{0}^{t}|\dot{\gamma}|=c t$. So $t$ is necessarily proportional to the arc length $s$.

In local coordinates, let $\gamma(t)=\left(x^{1}(t), \ldots, x^{n}(t)\right), \dot{\gamma}=\sum \dot{x}^{i}(t) \partial_{i}$. The geodesic equation is a second order non-linear ODE in $t$ :

$$
\nabla_{\dot{\gamma}} \dot{\gamma}=\sum_{k}\left(\ddot{x}^{k}+\sum_{i, j} \Gamma_{i j}^{k} \dot{x}^{i} \dot{x}^{j}\right) \partial_{k}=0 .
$$

Still, by the existence and uniqueness theorem, $\gamma$ exists and is uniquely determined by $\gamma\left(t_{0}\right)=p$ and $\gamma^{\prime}\left(t_{0}\right)=v$ for a maximal time interval.

The local existence ensures us to find the geodesics in any direction $v \in T_{p} M$. This leads to the notion of exponential map:

Definition 3.13 (Exponential map). For $p \in M$, the exponential map $\exp _{p}: U \subset T_{p} M \rightarrow M$ is defined by $\exp _{p}(t v)=\gamma(t)$, where $\gamma$ is the geodesic satisfying $\gamma(0)=p, \gamma^{\prime}(0)=v \in U$. Hence $\exp v=\gamma(1)$.


The exponential map $\exp _{p}$ is well-defined on some neighborhood $U$ of $0 \in T_{p} M$. But it may not be extended infinitely:

Example 3.14. Consider $M=\mathbb{R}^{2} \backslash\{0\}$. The geodesics are lines on the $\mathbb{R}^{2}$ plane. It is clear that the geodesic can not across the origin.


Proposition 3.15. $\exp _{p}$ is a local diffeomorphism near $0 \in T_{p} M$
The smoothness of $\exp _{p}$ follows from the smooth dependence of ODE on its initial values (cf. Section 8 of Chapter 1). By the inverse function theorem, we only need to compute

$$
d\left(\exp _{p}\right)_{0}: T_{0}\left(T_{p} M\right) \cong T_{p} M \rightarrow T_{p} M
$$

By direct calculation,

$$
d\left(\exp _{p}\right)_{0}(v)=\left.\frac{d}{d t}\right|_{t=0} \exp _{p}(t v)=v .
$$

Hence $d\left(\exp _{p}\right)_{0}=\mathrm{id}$ and the $C^{\infty}$ local inverse exists.

An easy application of the exponential map is the Riemann normal coordinates. Let $\left\{e_{1}, \ldots, e_{m}\right\}$ be an orthonormal basis of $T_{p} M$. In the diffeomorphic neighborhood $U$ of the exponential map, we define the Riemannian normal coordinates (RNC) by
$\exp _{p}\left(x^{1} e_{1}+x^{2} e_{2}+\ldots+x^{m} e_{m}\right) \mapsto\left(x^{1}, x^{2}, \ldots, x^{m}\right) \in U \subset \mathbb{R}^{m}$.


Why is the normal coordinates system good? An immediate consequence is the following striking fact:

Lemma 3.16. Under the normal coordinates, $\Gamma_{i j}^{k}(p)=0$ for all $i, j, k$.
Proof. Fix a direction $\mathbf{h}=\left(h^{1}, h^{2}, \ldots, h^{m}\right)$ and consider the line $\mathbf{x}(t)=\left(x^{1}(t), x^{2}(t), \ldots, x^{m}(t)\right)=\left(t h^{1}, t h^{2}, \ldots, t h^{m}\right)=t \mathbf{h}$ with corresponding geodesic $\gamma(t)=\exp _{p}(t \mathbf{h})$. The geodesic equation is

$$
\ddot{x}^{k}(t)+\sum_{i, j} \Gamma_{i j}^{k}(\gamma(t)) \dot{x}^{i}(t) \dot{x}^{j}(t)=0 .
$$

Since $\mathbf{x}(t)$ is linear in $t$, we get $\sum_{i, j} \Gamma_{i j}^{k}(\gamma(t)) h^{i} h^{j}=0$. Let $t=0$. Then for all $k$,

$$
\sum_{i, j} \Gamma_{i j}^{k}(p) h^{i} h^{j}=0
$$

Since this is true for arbitrary $\mathbf{h}$ in the existence neighborhood, we conclude that $\Gamma_{i j}^{k}(p)=0$.

As an application of the normal coordinates, we prove
Proposition 3.17. Geodesics are locally of minimal length.
More precisely, Take $B=B_{\rho}(0) \subset T_{p} M$ such that $\left.\exp _{p}\right|_{B}$ is a diffeomorphism. Any $q \in \exp _{p}(B)$ is connected to $p$ by a geodesic $\gamma$. Then $L(\gamma)$ is minimal among all piecewise smooth curves $\Gamma$ connecting $p$ and $q$.

Proof. Consider the spherical coordinate on $T_{p} M$ :

$$
\left(r, \theta^{1}, \ldots, \theta^{m-1}\right)
$$

which defines a "spherical normal coordinates" on $\exp _{p}(B)$. Denote $g_{r r}:=g\left(\partial_{r}, \partial_{r}\right), g_{r \alpha}:=g\left(\partial_{r}, \partial_{\theta^{\alpha}}\right), g_{\alpha \beta}:=g\left(\partial_{\theta^{\alpha}}, \partial_{\theta^{\beta}}\right)$.


Lemma 3.18 (Gauss lemma). $g_{r \alpha}=0$ for all $\alpha \in[1, m-1]$.
Assume the Gauss lemma, then

$$
d s^{2}=d r^{2}+\sum_{\alpha, \beta} g_{\alpha \beta} d \theta^{\alpha} \otimes d \theta^{\beta}
$$

Thus, the arc length of $\Gamma$ is

$$
\begin{aligned}
L(\Gamma) & =\int_{\Gamma} d s=\int \sqrt{\left(\frac{d r}{d t}\right)^{2}+\sum_{\alpha, \beta} g_{\alpha \beta} \frac{d \theta^{\alpha}}{d t} \frac{d \theta^{\beta}}{d t}} d t \\
& \geq \int\left|\frac{d r}{d t}\right| d t=L(\gamma)
\end{aligned}
$$

since $g$ is a positive definite symmetric tensor. The equality holds if and only if $\theta^{\alpha \prime}$ s are constants. That is, $\Gamma=\gamma$.

Proof of GaUss lemma. A simple proof using coordinates is left as an exercise (c.f. Exercise 3.5). The proof given below is intended to get us more familiar with the exponential map. Also we prove a stronger form: for $w \in T_{v}\left(T_{p} M\right) \cong T_{p} M$,

$$
\left\langle d\left(\exp _{p}\right)_{v}(v), d\left(\exp _{p}\right)_{v}(w)\right\rangle_{q}=\langle v, w\rangle_{p}
$$



Let $w=w_{1}+w_{2}$ where $w_{1} \perp v, w_{2}=\lambda v$. Then for the $w_{2}$ part,

$$
\left\langle d\left(\exp _{p}\right)_{v} v, d\left(\exp _{p}\right)_{v} w_{2}\right\rangle=\lambda\left\langle d\left(\exp _{p}\right)_{v} v, d\left(\exp _{p}\right)_{v} v\right\rangle=\lambda\langle v, v\rangle
$$

The last equality follows from the definition

$$
d\left(\exp _{p}\right)_{v} v=\left.\frac{d}{d t}\right|_{t=0} \exp _{p}(v+t v)=\left.\frac{d}{d t}\right|_{t=0} \gamma(t+1)=\gamma^{\prime}(1)
$$

and the fact that $\left|\gamma^{\prime}(t)\right|$ is constant in $t$. Hence $\left|\gamma^{\prime}(1)\right|=\left|\gamma^{\prime}(0)\right|=|v|$.
For the $w_{1}$ part, which is the Gauss lemma, we need to show

$$
\left\langle d\left(\exp _{p}\right)_{v} v, d\left(\exp _{p}\right)_{v} w_{1}\right\rangle=0
$$

Consider a parametrized surface in $M, \mathbb{X}(t, s):=\exp _{p}(t v(s))$, where $v(0)=v, v^{\prime}(0)=w_{1}$ and $|v(s)|$ is constant.


Denote $\mathbb{X}_{t}=\frac{\partial \mathbb{X}}{\partial t}=\mathbb{X}_{*} \frac{\partial}{\partial t}, \mathbb{X}_{s}=\frac{\partial \mathbb{X}}{\partial s}=\mathbb{X}_{*} \frac{\partial}{\partial s}$. Then

$$
\left\langle d\left(\exp _{p}\right)_{v} v, d\left(\exp _{p}\right)_{v} w_{1}\right\rangle_{q}=\left\langle\mathbb{X}_{t}, \mathbb{X}_{s}\right\rangle(1,0)
$$

By definition of geodesics,

$$
\begin{aligned}
\frac{\partial}{\partial t}\left\langle\mathbb{X}_{t}, \mathbb{X}_{s}\right\rangle & =\left\langle\nabla_{\partial_{t}} \mathbb{X}_{t}, \mathbb{X}_{s}\right\rangle+\left\langle\mathbb{X}_{t}, \nabla_{\partial_{t}} \mathbb{X}_{s}\right\rangle \\
& =\left\langle 0, \mathbb{X}_{s}\right\rangle+\left\langle\mathbb{X}_{t}, \nabla_{\partial_{s}} \mathbb{X}_{t}+\left[\mathbb{X}_{t}, \mathbb{X}_{s}\right]\right\rangle \\
& =\left\langle\mathbb{X}_{t}, \nabla_{\partial_{s}} \mathbb{X}_{t}\right\rangle=\frac{1}{2} \partial_{s}\left\langle\mathbb{X}_{t}, \mathbb{X}_{t}\right\rangle=0 .
\end{aligned}
$$

since $\left\langle\mathbb{X}_{t}, \mathbb{X}_{t}\right\rangle=|v(s)|^{2}$ is constant.

Exercise 3.5. Prove Gauss lemma by showing that $\Gamma_{r r}^{\alpha}=0$ and then $\partial_{r} g_{r \alpha}=0$ for all $\alpha$.
3. Metric space structure and completeness

A Riemannian manifold $(M, g)$ can be endowed with a metric space structure $(M, d)$ with $d(p, q)=\inf L(\gamma)$ among all curves $\gamma$ joining $p$ and $q$. From the Gauss lemma, we know that $B_{p}(r) \subset M$, the ball of radius $r$ centered at $p$, is precisely the diffeomorphic image of $B_{0}(r) \subset T_{p} M$ for small $r$. In particular the manifold topology coincides with the metric space topology.

Definition 3.19. A Riemannian manifold $(M, g)$ is complete if $(M, d)$ is a complete metric space.

In the definition of exponential map, $\exp _{p}$ is only defined on an open set $U$ of $0 \in T_{p} M$. It turns out that the question whether the map can be defined on whole $T_{p} M$ for every point $p \in M$ is equivalent to completeness of $(M, d)$. We first need the following

Lemma 3.20. Let $p \in M$, assume that $\exp _{p}$ is defined on whole $T_{p} M$, then any point $q \in M$ can be joined by a minimal geodesic from $p$.

Proof. Let $B_{p}(r)$ be the ball centered at $p$ in which $\exp _{p}$ is a diffeomorphism. If $q \in B_{p}(r)$, then we are done. Now, we consider $q \notin B_{p}(r)$, we claim that there exists $q^{\prime} \in \partial B_{p}(r)$ such that $d(p, q)=r+d\left(q^{\prime}, q\right)$. To see this, we easily have $d(p, q) \leq r+d\left(q^{\prime}, q\right)$ from triangle inequality. For the reverse inequality, let $q^{\prime} \in \partial B_{p}(r)$

such that $d\left(q^{\prime}, q\right)=\inf _{q^{\prime \prime} \in \partial B_{p}(r)} d\left(q^{\prime \prime}, q\right)$. By definition of our distance $d$, we get the reverse inequality.

Let $\gamma:[0, \infty) \rightarrow M$ be the geodesic enumerating from $p$ and passing through $q^{\prime}$. We next prove the statement by continuity method. Let $Z=\{t \in[0, \infty): d(p, \gamma(t))+d(\gamma(t), q)=d(p, q)\}$. First, $Z \neq \varnothing$ since $r \in Z$. Obviously, $Z$ is closed. Let $t_{0}=\sup Z$, we claim that $\gamma\left(t_{0}\right)=q$. Assume $\gamma\left(t_{0}\right) \neq q$. First, there exists $r_{1}>0$, consider $q^{\prime \prime} \in \partial B_{\gamma\left(t_{0}\right)}\left(r_{1}\right)$ such that

$$
d\left(\gamma\left(t_{0}\right), q^{\prime \prime}\right)+d\left(q^{\prime \prime}, q\right)=d\left(\gamma\left(t_{0}\right), q\right)
$$

where $d\left(\gamma\left(t_{0}\right), q^{\prime \prime}\right)=r_{1}$. Let $\sigma$ be the unique geodesic joining from $\gamma\left(t_{0}\right)$ to $q^{\prime \prime}$. Then from

$$
d\left(p, \gamma\left(t_{0}\right)+d\left(\gamma\left(t_{0}\right), q\right)=d(p, q)\right.
$$

we have:

$$
d\left(p, \gamma\left(t_{0}\right)\right)+d\left(\gamma\left(t_{0}\right), q^{\prime \prime}\right)=d\left(p, q^{\prime \prime}\right)
$$

However, this shows that $\left.\left.\gamma\right|_{\left[0, t_{0}\right]} \cup \sigma\right|_{\left[0, r_{1}\right]}$ is a minimal geodesic, which implies that $\gamma$ and $\sigma$ must fit together to form a smooth geodesic $\left.\gamma\right|_{\left[0, t_{0}+r_{1}\right]}$. Thus, $t_{0}+r_{1} \in Z$, a contradiction.

Theorem 3.21 (Hopf-Rinow-de Rham). The following are equivalent:
(1) $\exp _{p}$ is defined on all $T_{p} M$ for one $p$.
(2) $(M, g)$ is complete.
(3) $\exp _{p}$ is defined on all $T_{p} M$ for all $p \in M$.

Proof. (3) $\Rightarrow$ (1) is trivial. For (1) $\Rightarrow(2)$, let $\left\{q_{i}\right\}_{i=1}^{\infty}$ be a Cauchy sequence of $(M, d)$, and let $\gamma_{i}:\left[0, t_{i}\right] \rightarrow M$ be a sequence of minimal geodesic with $\gamma\left(t_{i}\right)=q_{i}$. Since $\gamma_{i}$ are parametrized by arclength, $\left\{t_{i}\right\}_{i=1}^{\infty}$ is also a Cauchy sequence in $[0, \infty)$, whose limit is denoted by $t_{0}$. Moreover, by compactness of $B_{1}(0) \subset T_{p} M$, we can pass
to a subsequence $\gamma_{i_{n}}$ such that $\gamma_{i_{n}}^{\prime} \rightarrow v \in T_{p} M$ as $n \rightarrow \infty$. Then consider the geodesic equation with initial value $\gamma(0)=p, \gamma^{\prime}(0)=v$, the Picard-Lindelöf theorem guarantees the continuity of solution of the ODE on initial data. As a result, we have $q_{i_{n}}=\gamma_{i_{n}}\left(t_{i_{n}}\right) \rightarrow \gamma\left(t_{0}\right)$, and thus $q_{i} \rightarrow \gamma\left(t_{0}\right)$ (since $q_{i}$ is Cauchy).

Next, for $(2) \Rightarrow(3)$, if there exists $p \in M$ such that a geodesic $\gamma$ with $\gamma(0)=p, \gamma^{\prime}(0)=v$ only defines on $\left[0, t_{0}\right)$. Pick any increasing sequence $t_{i} \nearrow t_{0}, \gamma\left(t_{i}\right)$ is a Cauchy sequence with limit $q$. Define $\gamma\left(t_{0}\right)=q$, then it is a continuous on $\left[0, t_{0}\right]$. Moreover, for sufficiently large $i, \gamma\left(t_{i}\right)$ lies in a normal coordinate ball of $q$, and thus we can join $\gamma\left(t_{i}\right)$ from $q$ by a minimal geodesic $\sigma$. Thus, the piecewise smooth curve $\gamma \cup \sigma$ must coincide together with a smooth geodesic, and $\gamma$ extends past $t_{0}$, a contradiction.

Corollary 3.22. That $(M, g)$ is complete implies it is geodesic convex, which means any two points can be joined by a minimal geodesic.

One should notice that even if $M$ is complete, two points may not be joined by a unique minimal geodesic, e.g. the north and the south poles of a sphere (cf. Section 6 for more on this).

At the end of this section, we prove Fact 2.37, namely the existence of convex neighborhood.


Given $p \in$ M. Fix $\delta>$ 0 such that $\forall q \in B_{p}(\delta)$, $\exp _{q}$ is a diffeomorphism on some $B_{q}\left(r_{q}\right)$. Simply let

$$
r_{0}=\min _{q \in \overline{B_{p}\left(\delta_{0}\right)}} r_{q}>0
$$

As in the picture, any $q, q^{\prime}$ may not be jointed by a geodesic inside the neighborhood.

Now, we regard $(M, g)$ as a metric space $(M, d)$. Reset $\delta=r_{0} / 4$. Let $W:=B_{p}\left(r_{0} / 4\right)$. Now for all $q \in W, \exp _{q}$ is a diffeomorphism on $W$ since $d\left(q, q^{\prime}\right) \leq d(q, p)+d\left(p, q^{\prime}\right)<r_{0} / 2$ for all $q, q^{\prime} \in W$. So there at least exists a minimal geodesic $\gamma$ connecting $q$ and $q^{\prime}$ with $\gamma \subset B_{q}\left(r_{0} / 2\right) \subset B_{p}\left(r_{0}\right)$.

We want to prove $\gamma \subset W=B_{p}\left(r_{0} / 4\right)$.


Let $\gamma(t)=\left(x^{1}(t), \ldots, x^{m}(t)\right)$. Recall the geodesic equation:

$$
\ddot{x}^{k}(t)+\Gamma_{i j}^{k} \dot{x}^{i}(t) \dot{x}^{j}(t)=0 .
$$

We will apply the maximum principle to it. Define the distance function $r(t):=d(p, \gamma(t)) . r^{2}=\sum_{k=1}^{m} x^{k} x^{k}, r r^{\prime}=\sum_{k=1}^{m} x^{k} \dot{x}^{k}$,

$$
\begin{aligned}
\left(r^{\prime}\right)^{2}+r r^{\prime \prime} & =\sum_{k} \ddot{x}^{k} x^{k}+\left(\dot{x}^{k}\right)^{2} \\
& =-\sum_{k} \Gamma_{i j}^{k} \dot{x}^{i} \dot{x}^{j} x^{k}+\sum_{k}\left(\dot{x}^{k}\right)^{2}
\end{aligned}
$$

Suppose $r$ reaches the max value at $t=t_{0}$. If $\gamma\left(t_{0}\right)=q$ or $q^{\prime}$ then we are done. Otherwise $r^{\prime}\left(t_{0}\right)=0, r^{\prime \prime}\left(t_{0}\right)<0$, hence LHS $\leq 0$. When $r_{0}$ is chosen small enough, RHS $>0$ which leads to a contradiction.
4. Riemann curvature tensor

Definition 3.23. The Riemann curvature detects the non-commutativity of covariant differentiations, that is:

$$
\begin{aligned}
R(X, Y) Z & :=\nabla_{\widetilde{X}} \nabla_{\widetilde{Y}} \widetilde{Z}-\nabla_{\widetilde{Y}} \nabla_{\widetilde{X}} \widetilde{Z}-\nabla_{[\widetilde{X}, \widetilde{Y}]} \widetilde{Z} \\
& \equiv\left(\left[\nabla_{\widetilde{X}^{\prime}} \nabla_{\widetilde{Y}}\right]-\nabla_{[\widetilde{X}, \widetilde{Y}]}\right) \widetilde{Z}
\end{aligned}
$$

for $X, Y, Z \in T_{p} M$ and $\widetilde{X}, \widetilde{Y}, \widetilde{Z}$ any vector fields that extend $X, Y, Z$.
The last term in definition is a correction to make it functionlinear. $R$ is a $(1,3)$ tensor. In local coordinates, we write

$$
R\left(\partial_{i}, \partial_{j}\right) \partial_{\ell}=: \sum R_{\ell i j}^{k} \partial_{k} .
$$

For convenience, we drop all summation symbols and use the convention that the appearance of an index $j$ on both upper and lower indices means summing over $j$. By computations,

$$
\begin{aligned}
R\left(\partial_{i}, \partial_{j}\right) \partial_{\ell} & =\nabla_{i} \nabla_{j} \partial_{\ell}-\nabla_{j} \nabla_{i} \partial_{\ell}-\nabla_{[i, j]} \partial_{\ell} \\
& =\nabla_{i}\left(\Gamma_{j \ell}^{s} \partial_{s}\right)-\nabla_{j}\left(\Gamma_{i \ell}^{t} \partial_{t}\right) \\
& =\left(\partial_{i} \Gamma_{j \ell}^{k}\right) \partial_{k}+\Gamma_{j \ell}^{s} \Gamma_{i s}^{k} \partial_{k}-\left(\partial_{j} \Gamma_{i \ell}^{k}\right) \partial_{k}-\Gamma_{i \ell}^{t} \Gamma_{j t}^{k} \partial_{k} \\
& =\left(\partial_{i} \Gamma_{j \ell}^{k}-\partial_{j} \Gamma_{i \ell}^{k}+\Gamma_{j \ell}^{s} \Gamma_{i s}^{k}-\Gamma_{i \ell}^{s} \Gamma_{j s}^{k}\right) \partial_{k} .
\end{aligned}
$$

Also, the metric $g$ leads to $T_{p} M \cong T_{p}^{*} M$ via $v \mapsto\langle v, \bullet\rangle$. In coordinates, this means we may use $g_{i j}$ to lowering the indices and define

$$
R_{k \ell i j}:=g_{k m} R_{\ell i j}^{m}=\left\langle R\left(\partial_{i}, \partial_{j}\right) \partial_{\ell}, \partial_{k}\right\rangle .
$$

Example 3.24. Let $Z=f^{p} \partial_{p}$ be a vector field. We denote

$$
\nabla_{i}\left(f^{p} \partial_{p}\right)=: f_{; i}^{p} \partial_{p}, \quad \nabla_{i} \nabla_{j}\left(f^{p} \partial_{p}\right)=: f_{; j i}^{p} \partial_{p}
$$

where indices after ";" means covariant differentiations. So

$$
\left(f_{; j i}^{p}-f_{; i j}^{p}\right) \partial_{p}=R\left(\partial_{i}, \partial_{j}\right) f^{p} \partial_{p}=f^{q} R_{q i j}^{p} \partial_{p}
$$

In other words, we can change the order of covariant differentiation via the commutation formula

$$
f_{; j i}^{p}=f_{; i j}^{p}+f^{q} R_{q i j}^{p}
$$

Exercise 3.6. Extend the covariant derivative to all tensors and prove the commutation formula:

$$
\begin{aligned}
T_{j_{1} \cdots j_{s} ; j i}^{i_{1} \cdots i_{r}}= & T_{j_{1} \cdots j_{s} ; i j}^{i_{1} \cdots i_{r}}+T_{j_{1} \cdots j_{s}}^{q i_{2} \cdots i_{r}} R_{q i j}^{i_{1}}+T_{j_{1} \cdots j_{s}}^{i_{1} q i_{3} \cdots i_{r}} R_{q j i}^{i_{2}}+\cdots \\
& +T_{p_{2} \cdots j_{s}}^{i_{1} \cdots i_{r}} R_{j_{1} j i}^{p}+\cdots+T_{j_{1} \cdots j_{s-1} p}^{i_{1} \cdots i_{r}} R_{j_{s} j i}^{p} .
\end{aligned}
$$

Proposition 3.25 (Symmetries on $R$ ).
(0) $R_{\ell i j}^{k}=-R_{\ell j i}^{k}$.
(1) Torsion free: the first Bianchi identity.

$$
R(X, Y) Z+R(Y, Z) X+R(Z, X) Y=0
$$

i.e. $R_{[\ell i j]}^{k}:=R_{\ell i j}^{k}+R_{i j \ell}^{k}+R_{j \ell i}^{k}=0$.
(2) Metrical: skew symmetry in $R_{k \ell i j}=-R_{\ell k i j}$.

Proof. (0) follow from $R(X, Y)=-R(Y, X)$. For (1), since these are all tensors, we may check it in any basis. Then

$$
\begin{aligned}
& R\left(\partial_{i}, \partial_{j}\right) \partial_{k}+R\left(\partial_{j}, \partial_{k}\right) \partial_{i}+R\left(\partial_{k}, \partial_{i}\right) \partial_{j} \\
& =\nabla_{i} \nabla_{j} \partial_{k}-\nabla_{j} \nabla_{i} \partial_{k}+\nabla_{j} \nabla_{k} \partial_{i}-\nabla_{k} \nabla_{j} \partial_{i}+\nabla_{k} \nabla_{i} \partial_{j}-\nabla_{i} \nabla_{k} \partial_{j}=0
\end{aligned}
$$

since $\nabla_{i} \partial_{j}-\nabla_{j} \partial_{i}=\left[\partial_{i}, \partial_{j}\right]=0$.
(2) is equivalent to $\langle R(X, Y) Z, W\rangle+\langle Z, R(X, Y) W\rangle=0$. Let $X=$ $\partial_{i}, Y=\partial_{j}$. Then

$$
\begin{aligned}
& \left\langle\nabla_{i} \nabla_{j} Z, W\right\rangle-\left\langle\nabla_{j} \nabla_{i} Z, W\right\rangle \\
& =\partial_{i}\left\langle\partial_{j} Z, W\right\rangle-\left\langle\nabla_{j} Z, \nabla_{i} W\right\rangle-\partial_{j}\left\langle\nabla_{i} Z, W\right\rangle+\left\langle\nabla_{i} Z, \nabla_{j} W\right\rangle \\
& =\partial_{i} \partial_{j}\langle Z, W\rangle-\partial_{i}\left\langle Z, \nabla_{j} W\right\rangle-\left\langle\nabla_{j} Z, \nabla_{i} W\right\rangle \\
& \quad \quad-\partial_{j} \partial_{i}\langle Z, W\rangle+\partial_{j}\left\langle Z, \nabla_{i} W\right\rangle+\left\langle\nabla_{i} Z, \nabla_{j} W\right\rangle,
\end{aligned}
$$

which is anti-symmetric in $Z, W$.
Exercise 3.7. Show that
(1) The symmetries in Proposition 3.25 implies $R_{k \ell i j}=R_{i j k \ell}$.
(2) The second Bianchi identity holds:

$$
0=R_{i j[k \ell ; m]}:=R_{i j k \ell ; m}+R_{i j \ell m ; k}+R_{i j m k ; \ell} .
$$

Definition 3.26 (Ricci and scalar curvature). The Ricci curvature tensor Ric $:=\sum R_{i j} d x^{i} \otimes d x^{j} \in C^{\infty}\left(\operatorname{Sym}^{2}\left(T^{*} M\right)\right)$ is the trace of $R$ :

$$
R_{i j}=\sum_{\ell} R_{i \ell j}^{\ell}=\sum_{k, \ell} g^{k \ell} R_{k i \ell j} .
$$

The further trace is a function called the scalar curvature:

$$
s:=\sum_{i, j} g^{i j} R_{i j} .
$$

Definition 3.27. A metric $g$ satisfying Ric $=\lambda g$ for some $\lambda \in C^{\infty}(M)$ is called an Einstein metric. ${ }^{3}$

[^2]Exercise 3.8. Let $(M, g)$ has $\operatorname{dim} M=m \geq 3$, if

$$
R_{i j}=\lambda g_{i j},
$$

then $\lambda$ is a constant. (Hint: use the Bianchi identity.)
There is another important curvature.
Definition 3.28 (Sectional curvature). For a two-dimensional plane $E \subset T_{p} M$, the sectional curvature is defined by

$$
K_{E} \equiv K(X, Y):=\frac{R(X, Y, X, Y)}{\|X \wedge Y\|^{2}}
$$

for any linearly independent vectors $X, Y \in E$.
In the two-dimensional case (surfaces), the sectional curvature is the only non-vanishing components in the Riemann curvature tensor. Moreover, it coincides with the Gaussian curvature:

Exercise 3.9. Let $X, Y \in T_{p} M$. Show that $K(X, Y)$ is the Gaussian curvature of the surface $\mathbb{X}(u, v)=\exp _{p}(u X+v Y)$.

Example 3.29. A sphere $S_{r}^{m}$ of radius $r$ has constant sectional curvature $K=1 / r^{2}$. The converse is true and will be proved later.

In fact, the Riemann curvature tensor is completely determined by the sectional curvature ranging over all two-planes.
where $T_{i j}$ is the energy momentum tensor in spacetime-a pseudo-Riemannian manifold $(M, g)$ where $g$ is non-degenerate but not necessarily positive definite. Most of our discussion on connection and curvature works in the general setup. (The spacetime considered by Einstein is 4-dimensional and is Lorentzian: it has local coordinates $\left(x^{0}=t, x^{1}, x^{2}, x^{3}\right)$, and $g$ has signature $(1,3)$.)

The spacetime is vacuum if $T_{i j} \equiv 0$. In this case,

$$
R_{i j}-\frac{s}{2} g_{i j}=0 \quad \Longrightarrow \quad g^{i j}\left(R_{i j}-\frac{s}{2} g_{i j}\right)=s-\frac{s}{2} m .
$$

If the spacetime dimension $m \neq 2$ then $s=0$, hence $R_{i j}=0$ as well.
For the uniform case, $T_{i j}=\mu g_{i j}$ for some function $\mu$, we have $R_{i j}=\lambda g_{i j}$ for $\lambda=\mu+s / 2$. This leads to the name of the Einstein metric.

Exercise 3.10 (Polarization formula). Let $\bar{K}(X, Y)=R(X, Y, X, Y)$. Then

$$
\begin{aligned}
\langle R(X, Y) Z, W\rangle= & \bar{K}(X+W, Y+Z)-\bar{K}(X+W, Y)-\bar{K}(X+W, Z) \\
& -\bar{K}(X, Y+Z)-\bar{K}(W, Y+Z)+\bar{K}(X, Z)+\bar{K}(W, Y) \\
& -\bar{K}(Y+W, X+Z)+\bar{K}(Y+W, X)+\bar{K}(Y+W, Z) \\
& +\bar{K}(Y, X+Z)+\bar{K}(W, X+Z)+\bar{K}(Y, Z)-\bar{K}(W, X) .
\end{aligned}
$$

Definition 3.30. We say that $(M, g)$ has constant sectional curvature $K(p)$ at $p$ if $K(X, Y)=K(p)$ for all independent vectors $X, Y \in T_{p} M$.

Exercise 3.11. $(M, g)$ has constant sectional curvature at $p$ if and only if $\langle R(X, Y) W, Z\rangle=K(p)(\langle X, Z\rangle\langle Y, W\rangle-\langle X, W\rangle\langle Y, Z\rangle)$. I.e.,

$$
R_{i j k \ell}=K(p)\left(g_{i k} g_{j \ell}-g_{i \ell} g_{j k}\right)
$$

## 5. Variations of geodesics

Let $\gamma:[a, b] \rightarrow M$ be (piecewise) $C^{\infty}$ curve on $M$. We consider the one parameter variation of $\gamma$ :

$$
F:[a, b] \times(-\epsilon, \epsilon) \rightarrow M
$$

with $F_{s}(t):=F(s, t), F_{0}=\gamma$. Let $T=F_{*} \frac{\partial}{\partial t}$ be the tangent vector field and $V=F_{*} \frac{\partial}{\partial s}$ be the variational vector field for $F_{s}(t)$. Then

$$
[V, T]=\left[F_{*} \frac{\partial}{\partial s}, F_{*} \frac{\partial}{\partial t}\right]=F_{*}\left[\frac{\partial}{\partial s}, \frac{\partial}{\partial t}\right]=0 .
$$

Exercise 3.12. Make sense of the above calculations as well as the formula $\nabla_{V} T=\nabla_{T} V$. Notice that in this special case we do not need to assume the vector fields are $F$-related.

The length of $F_{s}$ is given by the functional

$$
L(s)=\int_{a}^{b}\langle T, T\rangle^{1 / 2}(t, s) d t
$$




In this section, we will establish the minimality of arclength of geodesics from the variation method. We compute directly that

$$
\begin{align*}
L^{\prime}(s) & =\frac{\partial}{\partial s} \int_{a}^{b}\langle T, T\rangle^{1 / 2} d t=\int_{a}^{b} V\langle T, T\rangle^{1 / 2} d t  \tag{3.2}\\
& =\int_{a}^{b} \frac{2\left\langle\nabla_{V} T, T\right\rangle}{2\langle T, T\rangle^{1 / 2}} d t=\int_{a}^{b} \frac{\left\langle\nabla_{T} V, T\right\rangle}{\langle T, T\rangle^{1 / 2}} d t \quad \text { (torsion free) }  \tag{3.3}\\
& =\int_{a}^{b} \frac{T\langle V, T\rangle-V\left\langle\nabla_{T} T\right\rangle}{\langle T, T\rangle^{1 / 2}} d t . \quad \text { (metrical) } \tag{3.4}
\end{align*}
$$

We may assume that $\gamma$ is parametrized by its arc length, i.e. $\langle T, T\rangle(t, 0)=$ 1. Then we achieve the

Proposition 3.31 (First variation formula).

$$
\begin{equation*}
L^{\prime}(0)=\left.\langle V, T\rangle\right|_{a} ^{b}-\int_{a}^{b}\left\langle V, \nabla_{T} T\right\rangle d t \tag{3.5}
\end{equation*}
$$

Hence, $\gamma$ is a "critical point" of the arc-length functional $L$ if

$$
\left.\frac{d}{d s} L(s)\right|_{s=0}=0
$$

for any variation $F$ of $\gamma$ such that $F(a, s)=P, F(b, s)=Q$ with $P, Q \in M$ fixed. Note that $P, Q$ are fixed, so $V(a)=V(b)=0$. From the first variation formula we deduce

Fact 3.32. $\gamma$ is a critical point of $L$ among all end-points fixed variations if and only if $\gamma$ is a geodesic, i.e. $\nabla_{T} T=0$.

Remark 3.33. Instead of the ends-fixed condition, we may consider $\langle V, T\rangle(t, 0)=0$ for all $t \in[a, b]$, called "normal variations".

Assume $\gamma$ is parametrized by arc length. Differentiate (3.3) again

$$
\begin{aligned}
L^{\prime \prime}(s) & =\frac{\partial}{\partial s} \int_{a}^{b} \frac{\left\langle\nabla_{T} V, T\right\rangle}{\langle T, T\rangle^{1 / 2}} d t \\
& =\int_{a}^{b} \frac{V\left\langle\nabla_{T} V, T\right\rangle}{\langle T, T\rangle^{1 / 2}}-\frac{\left\langle\nabla_{T} V, T\right\rangle^{2}}{\langle T, T\rangle^{1 / 2}} d t .
\end{aligned}
$$

For the first term, we can change the order of differentiation by the curvature tensor:

$$
\begin{aligned}
V\left\langle\nabla_{T} V, T\right\rangle & =\left\langle\nabla_{V} \nabla_{T} V, T\right\rangle+\left\langle\nabla_{T} V, \nabla_{V} T\right\rangle \\
& =\left\langle\nabla_{T} \nabla_{V} V, T\right\rangle+\langle R(V, T) V, T\rangle+\left\|\nabla_{T} V\right\|^{2} \\
& =T\left\langle\nabla_{V} V, T\right\rangle-\left\langle\nabla_{V} V, \nabla_{T} T\right\rangle-\langle R(V, T) T, V\rangle+\left\|\nabla_{T} V\right\|^{2}
\end{aligned}
$$

Also, $\left\langle\nabla_{T} V, T\right\rangle=T\langle V, T\rangle-\left\langle V, \nabla_{T} T\right\rangle$. Hence we get the
Proposition 3.34 (Second variation formula-I). Let $\gamma:[a, b] \rightarrow M$ be a geodesic parametrized by arc length. Then for any variations,

$$
\begin{align*}
L^{\prime \prime}(0)= & \left.\left\langle\nabla_{V} V, T\right\rangle\right|_{a} ^{b} \\
& +\int_{a}^{b}\left(\left\|\nabla_{T} V\right\|^{2}-\langle R(V, T) T, V\rangle-(T\langle V, T\rangle)^{2}\right) d t \tag{3.6}
\end{align*}
$$

From

$$
\left\|\nabla_{T} V\right\|^{2}=\left\langle\nabla_{T} V, \nabla_{T} V\right\rangle=T\left\langle V, \nabla_{T} V\right\rangle-\left\langle V, \nabla_{T} \nabla_{T} V\right\rangle
$$

and

$$
\left\langle V, \nabla_{T} V\right\rangle=\left\langle V, \nabla_{V} T\right\rangle=V\langle V, T\rangle-\left\langle\nabla_{V} V, T\right\rangle
$$

we have the second form of this formula:
Proposition 3.35 (Second variation formula-II).

$$
\begin{align*}
L^{\prime \prime}(0)= & \left.V\langle V, T\rangle\right|_{a} ^{b}  \tag{3.7}\\
& -\int_{a}^{b}\left(\left\langle\nabla_{T}^{2} V+R(V, T) T, V\right\rangle+(T\langle V, T\rangle)^{2}\right) d t
\end{align*}
$$

Now we give some applications of these formulas. Recall that in a metric space $(M, d)$, its diameter is defined by

$$
\operatorname{diam} M:=\sup _{p, q \in M} d(p, q)
$$

Theorem 3.36 (Bonnet-Myers 1941). Let $\left(M^{m}, g\right)$ be complete with Ric $\geq$ $(m-1) K>0$ ( $K$ a constant). Then $M$ is compact with

$$
\operatorname{diam} M \leq \frac{\pi}{\sqrt{K}}
$$

PROOF. If there exist $p, q \in M$ with $d(p, q)>\frac{\pi}{\sqrt{K}}$, let $\gamma$ be a shortest (minimal) geodesic joining $p, q$, say $\gamma(0)=p, \gamma(\ell)=q$. Such a $\gamma$ exists by the Hopf-Rinow-de Rham theorem (Theorem 3.21).

Let $e_{1}=T=\gamma^{\prime}$ and pick $\left\{e_{i}(0)\right\}_{i=2}^{m}$ so that $\left\{e_{i}(0)\right\}_{i=1}^{m}$ is an O.N.B. of $T_{p} M$. Let $e_{i}(t)$ be the parallel translation of $e_{i}(0)$ along $\gamma$. Then $e_{i}(t)$ 's are still orthonormal for all $t$ by Lemma 3.10.

For each $i \in[2, m]$, let $V_{i}(t)=\sin (\pi t / \ell) e_{i}$ and construct the variation of $\gamma$ by

$$
F_{i}(t, s):=\exp _{\gamma(t)}\left(s V_{i}(t)\right)
$$

Since $F_{i}$ is an end-points fixed normal variation, the second variation formula in Proposition 3.6 implies that

$$
\begin{aligned}
0 & \leq \sum_{i=2}^{m} L_{i}^{\prime \prime}(0)=\sum_{i=2}^{m} \int_{0}^{\ell}\left(\left\|\nabla_{T} V_{i}\right\|^{2}-\left\langle R\left(V_{i}, T\right) T, V_{i}\right\rangle\right) d t \\
& =\sum_{i=2}^{m}\left(\int_{0}^{\ell}\left\|\frac{\pi}{\ell} \cos \left(\frac{\pi t}{\ell}\right) e_{i}\right\|^{2}-\sin ^{2}\left(\frac{\pi t}{\ell}\right)\left\langle R\left(e_{i}, e_{1}\right) e_{1}, e_{i}\right\rangle\right) d t
\end{aligned}
$$

Since $\left\|e_{1}\right\|=1$ and $\left.\operatorname{Ric}\left(e_{1}, e_{1}\right)=\sum_{m-2}^{m}\left\langle R\left(e_{i}, e_{1}\right) e_{1}, e_{i}\right)\right\rangle \geq(m-1) K$, a direct integration gives

$$
\sum_{i=2}^{m} L_{i}^{\prime \prime}(0) \leq \frac{m-1}{2}\left(\left(\frac{\pi}{\ell}\right)^{2}-K\right)<0
$$

a contradiction!

Remark 3.37. This was earlier proved by Bonnet in 1855 under the stronger assumption that the sectional curvature $\geq K>0$. The constant is sharp as shown by $S_{r}^{m}$. The converse was proved by S.-Y. Cheng in 1975: if there are $p, q \in M$ such that $d(p, q)=\pi / \sqrt{K}$ then $M$ is the round sphere with sectional curvature $K$.

Theorem 3.38 (Synge 1936). If $(M, g)$ is compact with sectional curvature $>0$, then $\pi_{1}(M)$ is finite. Moreover,
(1) If $m$ is even and $M$ is orientable, then $\pi_{1}(M)=\{1\}$.
(2) If $m$ is odd then $M$ must be orientable.

SKETCH OF PROOF. (1): if $\pi_{1}(M) \neq\{1\}$, then any non-trivial $\left[\gamma_{0}\right] \in \pi_{1}(M)$ can be represented by a $C^{\infty}$ closed curve $\gamma$ with shortest length. We leave the existence and smoothness of $\gamma$ as exercises.


Let $\gamma(0)=p$. For any $v=v(0) \in T_{p} M$, parallel translations of $v(0)$ along $\gamma$ gives $v(\ell) \in T_{p} M$. Parallel translations correspond to solving linear ODE. Hence the map $v(0) \mapsto v(\ell)=A v$ is a linear transformation on $T_{p} M$, called the holonomy along $\gamma$.

In general we know that $A \in O(m)$ since it preserves inner product. By our construction, $e:=\gamma^{\prime}(0)=\gamma^{\prime}(\ell)$ is invariant under $A$, hence we get the induced transformation

$$
\tilde{A}: e^{\perp} \rightarrow e^{\perp} \cong \mathbb{R}^{m-1}
$$

Now $m-1$ is odd and $M$ is orientable, we get $\tilde{A} \in S O(m-1)$ and it must has 1 as its eigenvalue. That is, $\tilde{A} v=v$ for some $v \in e^{\perp}$.

Let $v(t)$ be the parallel translation of $v=v(0)$ along $\gamma$ and use $V(t)=v(t)$ as the variation vector field. The second variation formula implies that

$$
\begin{aligned}
0 \leq L^{\prime \prime}(0) & =\int_{0}^{\ell}\left(\left\|\nabla_{T} V\right\|^{2}-\langle R(V, T) T, V\rangle\right) d t \\
& =-\int_{0}^{\ell} K(V, T) d t<0
\end{aligned}
$$

which is a contradiction. Thus $\pi_{1}(M)=\{1\}$.
We leave the proof of (2) as an exercise.

Exercise 3.13. Let $(M, g)$ be a compact Riemannian manifold. Show that any non-trivial $\left[\gamma_{0}\right] \in \pi_{1}(M)$ can be represented by a $C^{\infty}$ closed geodesic $\gamma \in\left[\gamma_{0}\right]$ of minimal length.

Exercise 3.14. Let $\left(M^{m}, g\right)$ be compact with positive sectional curvature. If $m$ is odd show that $M$ is orientable.
6. Jacobi fields

Recall the second variation formula of the form (Proposition 3.7)

$$
L^{\prime \prime}(0)=\left.V\langle V, T\rangle\right|_{a} ^{b}-\int_{a}^{b}\left\langle\nabla_{T}^{2} V+R(V, T) T, V\right\rangle+(T\langle V, T\rangle)^{2} d t
$$

In case of a normal variation, we found that $\nabla_{T}^{2} V+R(V, T) T$ is the essential quantity to determine this value.

Definition 3.39. For a geodesic $\gamma, T:=\gamma^{\prime}$, a vector field $V$ along $\gamma$ is a Jacobi field if it satisfies the Jacobi equation:

$$
\nabla_{T}^{2} V+R(V, T) T=0
$$

Since the Jacobi equation is a second order linear ODE, the Jacobi field $V$ is uniquely determined by $V(0)$ and $V^{\prime}(0)=\left.\nabla_{T} V\right|_{t=0}$. Also the solution space has dimension $2 m$.

Theorem 3.40. A vector field $J$ is Jacobi if and only if $J$ is a variation vector field of geodesics, i.e.

$$
J=\left.F_{*} \frac{\partial}{\partial s}\right|_{s=0}
$$

for some $F(t, s)$ which is a geodesic in for any s.


Proof. Assume that $F(t, s)$ is a family of geodesics for all $s$. Let $T=F_{*} \frac{\partial}{\partial t}, V=F_{*} \frac{\partial}{\partial s}$. Then

$$
\begin{aligned}
\nabla_{T}^{2} V & =\nabla_{T} \nabla_{T} V=\nabla_{T} \nabla_{V} T \\
& =\nabla_{V} \nabla_{T} T+R(T, V) T
\end{aligned}
$$

So we get the Jacobi equation since $\nabla_{T} T=0$ for all $s$.


Conversely, suppose that $J$ is a Jacobi field along a geodesic $\gamma$. Let $c(s)$ be a curve with $c(0)=p, c^{\prime}(0)=J(0)$ along $c$. Let $\check{T}, \check{J}^{\prime}$ be the parallel translation of $T(0), J^{\prime}(0)$. Define

$$
\begin{equation*}
F(t, s):=\exp _{c(s)}\left(t\left(\check{T}+s \check{J}^{\prime}\right)\right) \tag{3.8}
\end{equation*}
$$

which is clearly a variation of geodesics with $F(0, s)=c(s)$.
It suffice to check $V=F_{*} \frac{\partial}{\partial s}=J$, i.e. $V(0)=J(0), V^{\prime}(0)=J^{\prime}(0)$ :

$$
\begin{aligned}
V(0) & =\left.\left.\frac{\partial F}{\partial s}\right|_{(s, t)=(0,0)} \stackrel{(t=0)}{=} \frac{d}{d s} \exp _{c(s)}(0)\right|_{s=0}=c^{\prime}(0)=J(0) \\
V^{\prime}(0) & =\left.\nabla_{T} V\right|_{(s, t)=(0,0)}=\left.\nabla_{V} T\right|_{(s, t)=(0,0)} \\
& =\left.\nabla_{V}\left[\left(d \exp _{c(s)}\right)_{t\left(\tilde{T}+s \tilde{J}^{\prime}\right)} \frac{\partial}{\partial t}\left(t\left(\check{T}+s \check{J}^{\prime}\right)\right)\right]\right|_{(s, t)=(0,0)} \\
& \left.\stackrel{(t=0)}{=} \nabla_{\frac{\partial}{\partial s}} \operatorname{Id}\left(\check{T}+s \check{J}^{\prime}\right)\right|_{s=0}=\check{J}^{\prime}(0)=J^{\prime}(0) .
\end{aligned}
$$

The theorem is proved.
Example 3.41. On $\mathbb{S}^{2}$, it is clear that $\exp _{N}$ is well-defined on all $T_{N} S^{2}$. We observe that any Jacobi field $J$ with $J(0)=0$ must have $J(\pi)=0$. Indeed, any geodesic passing through the north pole $N$ is a great circle passing through the south pole $S$ as well. Hence the variation constructed above has $F(\pi, s)=S$ for all $s$.


Definition 3.42. For $P, Q \in \gamma$, we say $Q$ is a conjugate point of $P$ if $Q$ is a singular point of $\exp _{P}$, i.e. if $Q=\exp _{P}(\ell v)$ then

$$
\left(d \exp _{P}\right)_{\ell v}: T_{\ell v}\left(T_{P} M\right) \rightarrow T_{Q} M
$$

has a non-trivial kernel.


For any vector $w$ in the kernel, we will construct a Jacobi field with $J(0)=0$ and with $J(\ell)$ related to $w$.

Let $F(t, s):=\exp _{p}\left(t\left(v+s \frac{w}{\ell}\right)\right)$ (c.f. (3.8)). Then

$$
\left.\frac{\partial F}{\partial s}\right|_{s=0, t=\ell}=\left(d \exp _{p}\right)_{\ell v}(w)
$$

So if $J=\left.\frac{\partial F}{\partial s}\right|_{s=0}$ is Jacobi with $J(0)=0, J(\ell)=\left(d \exp _{p}\right)_{\ell v}(w)$ for general $w$. Hence $w$ is in the kernel if and only if $J(\ell)=0$ and $J^{\prime}(0)=$ $w / \ell$. And that is to say, $J(\ell), J^{\prime}(0)$ can determine each other if and only if $\left(d \exp _{p}\right)_{\ell v}$ is invertible. Thus, we have:

Corollary 3.43. $Q$ conjugates to $P$ if and only if there exists a Jacobi field $J \neq 0$ with $J(P)=J(Q)=0$. Furthermore, there's a correspondence
between the space of Jacobi fields and the kernel:

$$
\begin{aligned}
\left\{\begin{array}{c}
J J a c o b i, \\
J(P)=0=J(Q)
\end{array}\right\} & \leftrightarrow \operatorname{ker}\left(\left(\exp _{p}\right)_{*}\right)_{\ell v} \\
& J \mapsto w:=\ell J^{\prime}(0)
\end{aligned}
$$

Also, $Q$ is conjugate to $P$ if and only if $P$ is conjugate to $Q$.
Now we give some applications of Jacobi fields.
Theorem 3.44 (Hadamard 1898, Cartan 1928). If $M$ is complete and with $K \leq 0$, then $\forall p \in M, \exp _{p}: T_{p} M \rightarrow M$ is a covering map. In particular, the universal cover $\tilde{M}$ is diffeomorphic to $\mathbb{R}^{m}$.

Lemma 3.45. If $M$ complete with $K \leq 0$ then $\forall p \in M$, any geodesic $\gamma$ geodesic through $p$ does not have conjugate points of $p$ along $\gamma$.

Proof. Suppose there exists a conjugate point and let $J$ be a Jacobi field with $J(0)=J(\ell)=0$. By the Jacobi equation,

$$
\begin{aligned}
0 & =-\int_{0}^{\ell}\left\langle\nabla_{T}^{2} J+R(J, T) T, J\right\rangle d t \\
& =-\int_{0}^{\ell}\left(T\left\langle\nabla_{T} J, J\right\rangle-\left\|\nabla_{T} J\right\|^{2}+\langle R(J, T) T, J\rangle\right) d t \\
& =\left.\left\langle\nabla_{T} J, J\right\rangle\right|_{0} ^{\ell}+\int_{0}^{\ell}\left(\left\|\nabla_{T} J\right\|^{2}-\langle R(J, T) T, J\rangle\right) d t
\end{aligned}
$$

Since $J(0)=J(\ell)=0$ and $K \leq 0, \nabla_{T} J \equiv 0$ along $\gamma$. But this implies $J \equiv 0$, a contradiction.

Lemma 3.46. If $X$ is complete and $f: X \rightarrow Y$ is a local isometry, then $f$ is a covering map.

Exercise 3.15. Prove Lemma 3.46 and show by example that it fails if $X$ is not complete.

Proof of Cartan-HADAMARD THEOREM. By Lemma 3.45, $\exp _{p}$ is a local diffeomorphism on the whole $T_{p} M$. Hence

$$
(,):=\exp _{p}^{*}\langle,\rangle
$$

is a Riemannian metric on $T_{p} M \cong \mathbb{R}^{m}$ which makes $\exp _{p}$ a local isometry. In particular, $\exp _{0}: T_{0}\left(T_{p} M\right) \rightarrow T_{p} M$ is defined and is an isometry. By Hopf-Rinow theorem (cf. Theorem 3.21), ( $\left.T_{p} M,(),\right)$ is complete. By Lemma 3.46, $\exp _{p}$ is a covering map.
7. Local isometry and space forms

As the second application of Jacobi fields, we investigate how curvatures identifications are related to the constructions of isometries between spaces. In fact this is one of Riemann's original motivations to define his curvature tensor.

Let $X$ and $Y$ be Riemannian manifolds of the same dimension. For any $p \in X, q \in Y$, we may construct an isometry $\Phi: T_{p} X \rightarrow T_{q} Y$ of inner product spaces. Then $\Phi$ induces a local diffeomorphism:

$$
y=\varphi(x):=\exp _{q} \circ \Phi \circ \exp _{p}^{-1} .
$$

That is, we identify points $x \in X$ and $y \in Y$ by identifying the geodesics starting at $p$ and $q$ respectively.

Let $\gamma$ be a geodesic with $\gamma(0)=p$. Denote by $P_{\gamma}$ the parallel translation along $\gamma$. Also let $\tilde{\gamma}=\varphi \circ \gamma$ be the corresponding geodesic on $Y$. For $v \in T_{x} X$ where $x$ lies on $\gamma$, define

$$
g_{\gamma}: T_{x} X \rightarrow T_{\varphi(x)} Y, \quad v \mapsto g_{\gamma}(v):=P_{\tilde{\gamma}} \circ \Phi \circ P_{\gamma}^{-1}(v) .
$$

Then $g_{\gamma}$ is an isometry on inner product spaces.


Theorem 3.47 (Cartan1925, Ambrose 1956). If

$$
R\left(g_{\gamma}(v), g_{\gamma}(w)\right) g_{\gamma}(z)=g_{\gamma}(R(v, w) z)
$$

for all $v, w, z \in T_{x} X, x$ in a neighborhood of $p$, then $\varphi$ is a local isometry.

Proof. The key idea is to show that $g_{\gamma}$ is indeed the tangent map $(d \varphi)_{x}$, and then $\varphi$ is a local isometry.

Since $d \exp _{p}$ is local diffeomorphism, we may assume that there are no conjugate points in the neighborhood under our consideration. Let $J$ be the Jacobi field along $\gamma$ with $J(0)=0, J(\ell)=v$. Then

$$
\tilde{J}:=g_{\gamma}(J)
$$

is also a Jacobi field on $\tilde{\gamma}$ by the assumption.
Also, $\tilde{J}^{\prime}(0)=\Phi J^{\prime}(0)$ by the Chain Rule. Since $J(0)=0$, we have

$$
J(\ell)=\left(d \exp _{p}\right)_{\ell \gamma^{\prime}(0)} \ell J^{\prime}(0)
$$

as we had seen in the construction Jacobi fields (cf. (3.8)). Then

$$
\begin{aligned}
\tilde{J}(\ell) & =\left(d \exp _{q}\right)_{\ell \tilde{\gamma}^{\prime}(0)} \ell \tilde{J}^{\prime}(0) \\
& =\left(d \exp _{q}\right)_{\ell \tilde{\gamma}^{\prime}(0)} \circ \Phi \ell J^{\prime}(0) \\
& =\left(d \exp _{q}\right)_{\ell \tilde{\gamma}^{\prime}(0)} \circ \Phi \circ\left(d \exp _{p}\right)_{\ell \gamma^{\prime}(0)}^{-1} J(\ell) \\
& =d\left(\exp _{q} \circ \Phi \circ \exp _{p}^{-1}\right) J(\ell) \\
& =(d \varphi)_{x} J(\ell) .
\end{aligned}
$$

Here $d \Phi=\Phi$. This applies to all $v \in T_{x} X$, hence $(d \varphi)_{x}=g_{\gamma}$.
Ambrose indeed proved a global version of the theorem using piecewise geodesics, i.e. a continuous curve $\gamma=\bigcup_{i} \gamma_{i}$ which is a finite union of geodesics $\gamma_{i}$ (also called broken geodesics in the literature). Although we have not yet a map $\varphi: X \rightarrow Y$ as before, applying the exponential maps successively at end points of $\gamma_{i}$ 's the corresponding piecewise geodesic $\tilde{\gamma}$ in $Y$ is still defined. Then $g_{\gamma}$ can be defined similarly.

Theorem 3.48 (Cartan-Ambrose). Let X, Y be complete Riemannian with an isometry $\Phi: T_{p} X \rightarrow T_{q} Y$. Assume that $\pi_{1}(M)=\{1\}$. If

$$
R\left(g_{\gamma}(v), g_{\gamma}(w)\right) g_{\gamma}(z)=g_{\gamma}(R(v, w) z)
$$

for all piecewise geodesics $\gamma:[0, \ell] \rightarrow X$ with $\gamma(0)=p$ and $v, w, z \in$ $T_{\gamma(\ell)} X$, then for every other piecewise geodesic $\gamma_{1}$ with $\gamma_{1}(0)=p, \gamma_{1}(\ell)=$ $\gamma(\ell)$ we also have $\tilde{\gamma}_{1}(\ell)=\tilde{\gamma}(\ell)$.

This defines $\varphi: X \rightarrow Y$ which is a locally isometric covering map.

Exercise 3.16. Prove the global Cartan-Ambrose theorem.

Remark 3.49. Cartan proved his theorem for locally symmetric spaces and used it to classify them. Ambrose proved the general case as well as the global case. In 1959, Hicks generalized the result for general affine connections on the tangent bundle. Thus the theorem is often called the Cartan-Ambrose-Hicks theorem.

As model spaces, we often take manifolds with constant sectional curvature to start with. The more general model spaces including (locally) symmetric spaces require more tools to study and will be discussed in Chapter 5. Here are some simplest examples:

Example 3.50. Complete manifolds with constant sectional curvature.
(1) $\mathrm{S}^{n}, K>0$.
(2) $\mathbb{R}^{n} / \Gamma, \Gamma \cong \mathbb{Z}^{n}$. $K \equiv 0$.
(3) $\mathbb{H}^{n} / \Gamma, \Gamma$ is a discrete group of isometries. $K<0$. Here

$$
\mathbb{H}^{n}:=\left\{x \in \mathbb{R}^{n} \mid\|x\|<1\right\}, g=d s^{2}=\frac{4\|d x\|^{2}}{1-\|x\|^{2}}
$$

is known as the hyperbolic space.

Exercise 3.17. Show that $\mathbb{H}^{n}$ is complete with $K \equiv-1$.

Definition 3.51 (Space forms). A space form is a complete, simply connected Riemannian manifold with constant sectional curvature.

Now we will show that for each $K$, say $K \in\{-1,0,1\}$, the space forms are isometric:

Theorem 3.52. Let $(M, g)$ be a space form. Then
(1) $K \equiv-1 \Longleftrightarrow M \cong \mathbb{H}^{m}$.
(2) $K \equiv 0 \Longleftrightarrow M \cong \mathbb{R}^{m}$.
(3) $K \equiv 1 \Longleftrightarrow M \cong \mathbb{S}^{m}$.

Proof. For (1) and (2), $K \leq 0$. We compare $\mathbb{H}^{m}, \mathbb{R}^{m}$ with $M \ni p$. Step 1: Fix a linear isometry: $L: T_{0} \mathbb{H}^{m} \rightarrow T_{p} M$.

Step 2: Matching geodesics as in Cartan-Ambrose's theorem,

$$
\varphi=\exp _{p} \circ L \circ \exp _{0}^{-1}: U \ni 0 \rightarrow M
$$

Since $K \leq 0$, this map is well-defined by Cartan-Hadamard theorem (Theorem 3.44) and we need only geodesics.


In the case $K \equiv-1, \exp _{0}: T_{0} \mathbb{H}^{m} \rightarrow \mathbb{H}^{m}$ is a covering map. So $\varphi: \mathbb{H}^{m} \rightarrow M$ is a diffeomorphism. Cartan-Ambrose's theorem then implies that $\varphi$ is a local isometry.

The case $K=0$ is proved by the same argument. We leave the positive case as an exercise.

Exercise 3.18. Let $M$ be complete, simply connected with $K \equiv 1$. Show that $M \cong \mathbb{S}^{m}$.

So if $M$ has constant sectional curvature, $M \cong$ space form $/ \Gamma$ where $\Gamma$ is a discrete group of isometries, $\Gamma \cong \pi_{1}(M)$.

For example, we have hyperbolic geometry when $K=-1$.
Exercise 3.19. For the Poincaré upper half plane, show that all geodesic are circles that perpendicular to the real line. Use this to prove that the isometry group is precisely $S L(2, \mathbb{R})$ under the action

$$
\left[\begin{array}{ll}
a & b \\
c & d
\end{array}\right](z)=\frac{a z+b}{c z+d} .
$$


8. Variations of higher dimensional submanifolds

In this final section we consider the variational properties on submanifolds. Let

$$
f: M^{m} \hookrightarrow \bar{M}^{m+k}
$$

be a Riemannian imbedding. That is, $(\bar{M}, \bar{g})$ is equipped with the Levi-Civita connection $\bar{\nabla}=\bar{\nabla}^{L C}$, and $(M, g)$ is with the induced metric $g=f^{*} \bar{g}$ and hence $\nabla:=\nabla^{L C}=\bar{\nabla}^{T}$, where $T$ means the tangential projection under $T_{p} \bar{M}=T_{p} M \oplus^{\perp} N_{p}$.


In general, for $X, Y, Z, W \in T_{p} M$,

$$
\langle R(X, Y) Z, W\rangle \neq\langle\bar{R}(X, Y) Z, W\rangle
$$

E.g. for $\mathbb{S}^{m} \hookrightarrow \mathbb{R}^{m+1}, K \equiv 1$ on the $S^{m}$ while $K \equiv 0$ on $\mathbb{R}^{m}$.

In order to investigate the curvature defect, we start with the defect of covariant differentiations:

Definition 3.53 (Second fundamental form). The second fundamental form $B \in C^{\infty}\left(M, \operatorname{Sym}^{2}(N)\right)$ is a symmetric bilinear form with values
in the normal bundle: for $X, Y \in T_{p} M$,

$$
\mathbf{I I}(X, Y) \equiv B(X, Y):=\left(\bar{\nabla}_{X} \tilde{Y}\right)^{N}
$$

Here $\tilde{Y}$ is any vector field on $\bar{M}$ extending $Y$, and $N$ is the normal projection to $N_{p}$. The trace of $B$ is called the mean curvature (vector):

$$
\vec{H}=H=\operatorname{tr}(B) \in N
$$

Explanations: $B$ is tensorial in $X$. It is symmetric since

$$
B(X, Y)-B(Y, X)=\left(\bar{\nabla}_{X} \tilde{Y}-\bar{\nabla}_{Y} \tilde{X}\right)^{N}=[\tilde{X}, \tilde{Y}]^{N}=0,
$$

hence $B$ is also tensorial in $Y$. The trace of $B$ can be computed by

$$
H(p)=\sum_{i=1}^{m} B\left(e_{i}, e_{j}\right)=\sum_{i, j=1}^{m} g^{i j} B\left(\partial_{i}, \partial_{j}\right),
$$

where $e_{1}, \ldots, e_{m}$ is any O.N.B. of $T_{p} M$ and $\partial_{i}=\partial / \partial x^{i}$ are coordinate vector fields in any coordinate system.

The relation between $R$ and $\bar{R}$ is described by
Proposition 3.54 (Gauss equation). For any $X, Y, Z, W \in T_{p} M$,

$$
\begin{align*}
\langle R(X, Y) W, Z\rangle= & \langle\bar{R}(X, Y) W, Z\rangle \\
& +\langle B(X, Z), B(Y, W)\rangle-\langle B(X, W), B(Y, Z)\rangle \tag{3.9}
\end{align*}
$$

Proof. Denote the extensions of a vector by the same symbol. Direct computation gives

$$
\begin{aligned}
\langle & R(X, Y) W, Z\rangle=\left\langle\nabla_{X} \nabla_{Y} W, Z\right\rangle-\left\langle\nabla_{Y} \nabla_{X} W, Z\right\rangle-\left\langle\nabla_{[X, Y]} W, Z\right\rangle \\
= & \left\langle\bar{\nabla}_{X} \nabla_{Y} W, Z\right\rangle-\left\langle\bar{\nabla}_{Y} \nabla_{X} W, Z\right\rangle-\left\langle\bar{\nabla}_{[X, Y]} W, Z\right\rangle \\
= & X\left\langle\bar{\nabla}_{Y} W, Z\right\rangle-\left\langle\nabla_{Y} W, \bar{\nabla}_{X} Z\right\rangle-Y\left\langle\bar{\nabla}_{X} W, Z\right\rangle+\left\langle\nabla_{X} W, \bar{\nabla}_{Y} Z\right\rangle \\
& \quad-\left\langle\bar{\nabla}_{[X, Y]} W, Z\right\rangle \\
= & X\left\langle\bar{\nabla}_{Y} W, Z\right\rangle-\left\langle\bar{\nabla}_{Y} W, \bar{\nabla}_{X} Z\right\rangle+\left\langle B(Y, W), \bar{\nabla}_{X} Z\right\rangle \\
& \quad-Y\left\langle\bar{\nabla}_{X} W, Z\right\rangle+\left\langle\bar{\nabla}_{X} W, \bar{\nabla}_{Y} Z\right\rangle-\left\langle B(X, W), \bar{\nabla}_{Y} Z\right\rangle \\
& \quad-\left\langle\bar{\nabla}_{[X, Y]} W, Z\right\rangle \\
= & \langle\bar{R}(X, Y) W, Z\rangle+\langle B(X, Z), B(Y, W)\rangle-\langle B(X, W), B(Y, Z)\rangle .
\end{aligned}
$$

Notice that $T \bar{M}=T M \oplus^{\perp} N$ is used in most steps.

Example 3.55. Let $M^{m} \hookrightarrow \mathbb{R}^{m+1}$ be a hypersurface with prescribed unit normal vector field $N$. Choose a local coordinate $\left(x^{1}, \ldots, x^{m}\right)$. Denote $\partial_{i}=\mathbb{X}_{i}=\frac{\partial}{\partial x^{i}}$ and

$$
h_{i j} N:=B\left(\partial_{i}, \partial_{j}\right)=\left(\nabla_{\partial_{i}} \partial_{j}\right)^{N}=\left(\partial_{i} \mathbb{X}_{j}\right)^{N}=\mathbb{X}_{j i}^{N}
$$

Then the Gauss equation reads as

$$
R_{i j k \ell}=\bar{R}_{i j k \ell}+h_{i k} h_{j \ell}-h_{i \ell} h_{j k} .
$$

E.g., for $M=\mathbb{S}_{r}^{m} \hookrightarrow \mathbb{R}^{m+1}$ with normal vector $N=\mathbb{X} / r$,

$$
h_{i j}=\mathbb{X}_{i j} \cdot N=\left(\mathbb{X}_{i} \cdot N\right)_{j}-\left(\mathbb{X}_{i} \cdot N_{j}\right)=-\frac{1}{r} \mathbb{X}_{i} \cdot \mathbb{X}_{j}=-\frac{1}{r} g_{i j} .
$$

Hence $R_{i j k \ell}=\frac{1}{r^{2}}\left(g_{i k} g_{j \ell}-g_{i \ell} g_{j k}\right)$.
Since an immersion is locally an imbedding, and all of the above discussions are local, hence they works for immersions with small modifications.

Given an immersion $\imath: M^{n} \rightarrow\left(N^{n+p}, \bar{g}\right)$ with the induced metric on $(M, g)$, we consider the variation of $M$ in $N$ :

$$
F:(-\epsilon, \epsilon) \times M \longrightarrow N,
$$

which is $C^{\infty}$ with $F_{0}=\imath, F_{s}=F(s, \bullet): M \rightarrow N . F_{s}$ is still an immersion for $s$ small. Let $\eta:=F_{*} \frac{\partial}{\partial s}$ be the variation vector field.

For simplicity, we assume that $M$ is oriented. Nevertheless we allow $M$ to have non-empty boundary $\partial M$ with outer normal $v$.

Recall the $n$-dimensional area functional:

$$
A(s)=\int_{M_{s}} d A_{s}=\int_{M} \sqrt{g(s)} d x^{1} \wedge \cdots \wedge d x^{n}
$$

where $g_{s}=F_{s}^{*} \bar{g}$ and $g(s)=\operatorname{det}\left(g_{s, i j}\right)$. Taking derivative in $s$ :

$$
A^{\prime}(s)=\int_{M} \frac{\partial}{\partial s} \sqrt{g(s)} d x^{1} \wedge \ldots \wedge d x^{n}
$$

Let $G_{s, i j}$ be the cofactor of $g_{s, i j}$. To save notations, we omit the subscript $s$ in the presentation when no confusion will arise. Then

$$
g^{\prime}(s)=g_{i j}^{\prime} G_{i j}=g_{i j}^{\prime} g^{i j} g .
$$

Hence

$$
\begin{equation*}
A^{\prime}(s)=\int_{M} \frac{g_{i j}^{\prime} g^{i j} g}{2 \sqrt{g}} d x^{1} \wedge \cdots \wedge d x^{n}=\frac{1}{2} \int_{M} g_{i j}^{\prime} g^{i j} d A_{s} \tag{3.10}
\end{equation*}
$$

Theorem 3.56 (First variation formula). For a variation of an immersed submanifold $M \leftrightarrow N$ with variation field $\eta$ and outer normal field $v$,

$$
A^{\prime}(0)=\int_{\partial M}\left\langle\eta^{T}, v\right\rangle d \sigma-\int_{M}\langle\eta, \vec{H}\rangle d A
$$

Proof. To calculate $A^{\prime}(0)$, we will determine the integrand tensor in (3.10) at each point. Let $p \in M$. Pick the Riemann normal coordinate $\left(x^{1}, \ldots, x^{n}\right)$ of $\left(M, g=g_{0}\right)$ at $p$. Then

$$
g_{i j}^{\prime}(0)=\left.\frac{\partial}{\partial s}\left\langle\partial_{i}, \partial_{j}\right\rangle\right|_{s=0}=\left\langle\bar{\nabla}_{\eta} \partial_{i}, \partial_{j}\right\rangle+\left\langle\partial_{i}, \bar{\nabla}_{\eta} \partial_{j}\right\rangle .
$$

Since $\left(\nabla_{\partial_{i}} \partial_{j}\right)(p)=\left.\sum_{k} \Gamma_{i j}^{k}(p) \partial_{k}\right|_{p}=0$ by Lemma 3.16, we have

$$
\begin{aligned}
\left\langle\bar{\nabla}_{\partial_{i}} \eta, \partial_{j}\right\rangle & =\partial_{i}\left\langle\eta, \partial_{j}\right\rangle-\left\langle\eta, \bar{\nabla}_{\partial_{i}} \partial_{j}\right\rangle \\
& =\partial_{i}\left\langle\eta, \partial_{j}\right\rangle-\left\langle\eta, \nabla_{\partial_{i}} \partial_{j}+B\left(\partial_{i}, \partial_{j}\right)\right\rangle \\
& =\partial_{i}\left\langle\eta, \partial_{j}\right\rangle-\left\langle\eta, B\left(\partial_{i}, \partial_{j}\right)\right\rangle .
\end{aligned}
$$

The other term $\left\langle\partial_{i}, \bar{\nabla}_{\eta} \partial_{j}\right\rangle$ is obtained by symmetry, and after contracting with $g^{i j}$ they are equal. Hence

$$
\begin{aligned}
A^{\prime}(0) & =\int_{M} g^{i j}\left(\partial_{i}\left\langle\eta, \partial_{j}\right\rangle-\left\langle\eta, B\left(\partial_{i}, \partial_{j}\right)\right) d A\right. \\
& =\int_{M} \operatorname{div} \eta^{T} d A-\int_{M}\langle\eta, \vec{H}\rangle d A \\
& =\int_{\partial M}\left\langle\eta^{T}, v\right\rangle d \sigma-\int_{M}\langle\eta, \vec{H}\rangle d A
\end{aligned}
$$

where the last equality follows from the Stokes theorem.
It is clear that the critical point condition $A^{\prime}(0)=0$ for all $\eta$ which vanishes on $\partial M$ is equivalent to $\vec{H}=0$. Following the tradition, these "critical immersed submanifolds" are called "minimal":

Definition 3.57. $M$ is a (immersed) minimal submanifold in $N$ if $\vec{H}=0$.

Remark 3.58. One would ask if a minimal submanifold is indeed "locally minimal" in the usual sense. One would also ask about the existence problem. E.g., given a piecewise smooth curve $\Gamma$ in $\mathbb{R}^{3}$, is there an area minimizing surface $S$ with $\partial S=\Gamma$ ? We will discuss these questions in later chapters.

We proceed to study the second variations. Differentiating (3.10) again we get

$$
\begin{aligned}
A^{\prime \prime}(s) & =\frac{1}{2} \int_{M} \frac{\partial}{\partial s}\left(g_{i j}^{\prime} g^{i j} \sqrt{g}\right) d x^{1} \wedge \ldots \wedge d x^{n} \\
& =\int_{M}(\underbrace{\frac{1}{2} g_{i j}^{\prime \prime} g^{i j} \sqrt{g}}_{(1)}+\underbrace{\frac{1}{2} g_{i j}^{\prime}\left(g^{i j}\right)^{\prime} \sqrt{g}}_{(2)}+\underbrace{\frac{1}{2} g_{i j}^{\prime} g^{i j}(\sqrt{g})^{\prime}}_{(3)}) d x^{1} \wedge \cdots \wedge d x^{n}
\end{aligned}
$$

At $p \in M, s=0$, we pick the Riemann normal coordinate so that $g_{i j}(p)=\delta_{i j},\left(\nabla_{\partial_{i}} \partial_{j}\right)(p)=0$. Then

$$
\begin{aligned}
(1)= & \eta\left\langle\bar{\nabla}_{\partial_{i}} \partial_{i}, \partial_{i}\right\rangle=\left\langle\bar{\nabla}_{\eta} \bar{\nabla}_{\eta} \partial_{i}, \partial_{i}\right\rangle+\left\langle\bar{\nabla}_{\eta} \partial_{i}, \bar{\nabla}_{\eta} \partial_{i}\right\rangle \\
= & \left\langle\bar{\nabla} \bar{\nabla}_{\partial_{i}} \eta, \partial_{i}\right\rangle+\left\|\bar{\nabla}_{\partial_{i}} \eta\right\|^{2} \\
= & \left\langle\bar{R}\left(\eta, \partial_{i}\right) \eta, \partial_{i}\right\rangle+\left\langle\bar{\nabla}_{\partial_{i}} \bar{\nabla}_{\eta} \eta, \partial_{i}\right\rangle+\left\|\bar{\nabla}_{\partial_{i}} \eta^{T}\right\|^{2}+\left\|\bar{\nabla}_{\partial_{i}} \eta^{N}\right\|^{2} \\
= & -\bar{R}\left(\eta, \partial_{i}, \eta, \partial_{i}\right)+\partial_{i}\left\langle\bar{\nabla}_{\eta} \eta, \partial_{i}\right\rangle-\left\langle\bar{\nabla}_{\eta} \eta, \bar{\nabla}_{\partial_{i}} \partial_{i}\right\rangle \\
& \quad+\left\|\bar{\nabla}_{\partial_{i}} \eta^{T}\right\|^{2}+\left\|\bar{\nabla}_{\partial_{i}} \eta^{N}\right\|^{2} .
\end{aligned}
$$

For (2), recall that $\left(G^{-1}\right)^{\prime}=-G^{-1} G^{\prime} G^{-1}$. Hence

$$
\begin{aligned}
(2) & =-\frac{1}{2} \sum_{i, j} g_{i j}^{\prime}\left(g^{i k} g_{k l}^{\prime} g^{\ell j}\right)=-\frac{1}{2} \sum_{i, j}^{n}\left(g_{i j}^{\prime}\right)^{2} \\
& =-\frac{1}{2} \sum_{i, j}\left(\left\langle\bar{\nabla}_{\eta} \partial_{i}, \partial_{j}\right\rangle+\left\langle\partial_{i}, \bar{\nabla}_{\eta} \partial_{j}\right)^{2}\right.
\end{aligned}
$$

For (3), we have seen that $\sqrt{g}^{\prime}=\frac{1}{2} \sum g_{i j}^{\prime} g^{i j}$, hence

$$
(3)=\frac{1}{4}\left(\sum_{i, j} g_{i j}^{\prime} g^{i j}\right)^{2}=\left(\sum\left\langle\bar{\nabla}_{\eta} \partial_{i}, \partial_{i}\right\rangle\right)^{2} .
$$

By putting together (1), (2) and (3) we get

Theorem 3.59 (Second variation formula). Under the Riemann normal coordinates,

$$
\begin{aligned}
A^{\prime \prime}(0)= & \int_{\partial M}\left\langle\bar{\nabla}_{\eta} \eta, \nu\right\rangle-\int_{M}\left(\left\langle\bar{\nabla}_{\eta} \eta, \vec{H}\right\rangle+\sum_{i}\left\langle\bar{R}\left(\partial_{i}, \eta\right) \eta, \partial_{i}\right\rangle\right) \\
& +\int_{M} \sum_{i, j}\left\langle\bar{\nabla}_{\partial_{i}} \eta, \partial_{j}\right\rangle^{2}+\sum_{i}\left\|\bar{\nabla}_{\partial_{i}} \eta^{N}\right\|^{2} \\
& -\frac{1}{2} \int_{M} \sum_{i, j=1}^{n}\left(\left\langle\bar{\nabla}_{\eta} \partial_{i}, \partial_{j}\right\rangle+\left\langle\partial_{i}, \bar{\nabla}_{\eta} \partial_{j}\right\rangle\right)^{2} \\
& +\int_{M}\left(\sum_{i}\left\langle\bar{\nabla}_{\eta} \partial_{i}, \partial_{i}\right\rangle\right)^{2} .
\end{aligned}
$$

The expression is complicate. Nevertheless, in many specific applications it can be simplified considerably.

For $i: M \hookrightarrow N$ (minimal), $F$ normal variation ( $\eta \perp T M$ ), let $\eta=f \mu$ with $\mu$ being an unit normal vector field. In this case,

$$
\begin{aligned}
\left\langle\bar{\nabla}_{\eta} \partial_{i}, \partial_{j}\right\rangle & =\left\langle\bar{\nabla}_{\partial_{i}} \eta, \partial_{j}\right\rangle=\partial_{i}\left\langle\eta, \partial_{j}\right\rangle-\left\langle\eta, \bar{\nabla}_{\partial_{i}} \partial_{j}\right\rangle \\
& =-\left\langle\eta, B\left(\partial_{i}, \partial_{j}\right)\right\rangle
\end{aligned}
$$

and the calculation for $\left\langle\partial_{i}, \bar{\nabla}_{\eta} \partial_{j}\right\rangle$ is the same. So

$$
-\frac{1}{2} \int_{M} \sum_{i, j=1}^{n}\left(\left\langle\bar{\nabla}_{\eta} \partial_{i}, \partial_{j}\right\rangle+\left\langle\partial_{i}, \bar{\nabla}_{\eta} \partial_{j}\right\rangle\right)^{2}=-2 \int_{M} \sum_{i, j=1}^{n}|f|^{2}\left\langle\mu, B\left(\partial_{i}, \partial_{j}\right)\right\rangle^{2}
$$

Combining with the term $\sum\left\langle\bar{\nabla}_{\partial_{i}} \eta, \partial_{j}\right\rangle^{2}$, we get

$$
-\int_{M}|f|^{2}\|\langle B, \mu\rangle\|^{2}
$$

Also,

$$
\sum_{i}\left\langle\bar{\nabla}_{\eta} \partial_{i}, \partial_{i}\right\rangle=\sum_{i}\left\langle\bar{\nabla}_{\partial_{i}} \eta, \partial_{i}\right\rangle=-f \sum_{i}\left\langle\bar{\nabla}_{\partial_{i}} \partial_{i}, \mu\right\rangle=-f\langle\vec{H}, \mu\rangle .
$$

For $\sum_{i}\left\|\bar{\nabla}_{\partial_{i}}^{N} \eta\right\|^{2}$, if $M$ is a hypersurface we may write $\eta=f e_{n+1}$ ( $v=e_{n+1}$ ), and then

$$
\bar{\nabla}_{\partial_{i}} \eta=\bar{\nabla}_{\partial_{i}}\left(f e_{n+1}\right)=\left(\partial_{i} f\right) e_{n+1}+f \bar{\nabla}_{\partial_{i}} e_{n+1}
$$

Since $\left\langle\bar{\nabla}_{\partial_{i}} e_{n+1}, e_{n+1}\right\rangle=\frac{1}{2} \partial_{i}\left\langle e_{n+1}, e_{n+1}\right\rangle=0, f \bar{\nabla}_{\partial_{i}} e_{n+1} \in T M$. So

$$
\bar{\nabla}_{\partial_{i}}^{N} \eta=\left(\partial_{i} f\right) e_{n+1}, \quad \sum_{i}\left\|\bar{\nabla}_{\partial_{i}}^{N} \eta\right\|^{2}=\|d f\|^{2}=\|\nabla f\|^{2}
$$

We have thus a corollary:
Corollary 3.60. For $\partial M$-fixed normal variations of a minimal immersion $M \rightarrow N$,

$$
A^{\prime \prime}(0)=\int_{M}\left(\|\nabla f\|^{2}-\left(\overline{\operatorname{Ric}}(\vec{n}, \vec{n})+\|B\|^{2}\right) f^{2}\right)
$$

where $\vec{n}=e_{n+1}$.
As an application, we prove a theorem due to Schoen and Yau which was used in their 1979 proof of the "positive mass theorem" in general relativity.

Definition 3.61. A minimal immersion $M \rightarrow N$ is stable if $A^{\prime \prime}(0) \geq 0$ for all variations which fixed $\partial M$.

Theorem 3.62 (Schoen-Yau). Let $\bar{M}$ is a compact oriented 3-dimensional manifold with non-negative scalar curvature and $\bar{s}(p)>0$ for some $p \in$ $M$. Then there is no stable minimal immersed surface $M \rightarrow N$ which is compact orientable, $\partial M=\varnothing$, and $g(M) \geq 1$.

Proof. If $M$ is stable, i.e. $A^{\prime \prime}(0) \geq 0$, then for all $f \in C^{\infty}(M)$ :

$$
\int_{M}\left(\overline{\operatorname{Ric}}(\vec{n}, \vec{n})+\|B\|^{2}\right) f^{2} \leq \int_{M}|\nabla f|^{2}
$$

Let $f \equiv 1$ and apply the follwoing lemma:
Lemma 3.63. Let $M^{2} \rightarrow \bar{M}^{3}$ be a minimal immersion. Then

$$
\overline{\operatorname{Ric}}(\vec{n}, \vec{n})=\frac{1}{2} \bar{s}-K-\frac{1}{2}\|B\|^{2}
$$

where $K$ is the Gauss curvature of $M$.
With this lemma, we then have

$$
\int_{M}\left(\frac{1}{2} \bar{s}+\frac{1}{2}\|B\|^{2}\right)-\int_{M} K \leq 0
$$

By the Gauss-Bonnet theorem, we have

$$
\int_{M} K=2 \pi \chi(M)=4 \pi(1-g(M)) \leq 0
$$

if $g(M) \geq 1$. This contradicts to the assumption on $\bar{s}$.

Proof of Lemma. Pick Riemann normal coordinate at $p$. Then

$$
\begin{aligned}
& \qquad \begin{array}{l}
\bar{R}=\bar{g}^{i j} \bar{g} k \ell \bar{R}_{i k j \ell}=2\left(\bar{R}_{1212}+\bar{R}_{1313}+\bar{R}_{2323}\right) \\
=2\left(\overline{\operatorname{Ric}}\left(e_{3}, e_{3}\right)+\bar{R}_{1212}\right) .
\end{array} \\
& \text { By Gauss' equation } R_{i j k \ell}=\bar{R}_{i j k \ell}+h_{i k} h_{j \ell}-h_{i \ell} h_{j k} \text {, we get } \\
& K=R_{1212}=\bar{R}_{1212}+h_{11} h_{22}-h_{12} h_{21} .
\end{aligned}
$$

Since $M$ is minimal, we have

$$
B=\left[\begin{array}{ll}
h_{11} & h_{12} \\
h_{21} & h_{22}
\end{array}\right], \quad \operatorname{tr}(B)=h_{11}+h_{22}=0
$$

Hence $\|B\|^{2}=h_{11}^{2}+h_{12}^{2}+h_{21}^{2}+h_{22}^{2}=2 h_{11}^{2}+2 h_{12}^{2}$ and $h_{11} h_{22}-$ $h_{12} h_{21}=-\left(h_{11}^{2}+h_{12}^{2}\right)=\frac{1}{2}\|B\|^{2}$. The formula follows.
9. Problems

1. Determine all the geodesics on the sphere $S^{n}$.
2. ([Car92] Ch.2 \#8) Consider the upper half-plane

$$
\mathbb{R}_{+}^{2}=\left\{(x, y) \in \mathbb{R}^{2} \mid y>0\right\}
$$

with the metric given by $g_{11}=g_{22}=\frac{1}{y^{2}}, g_{12}=0$ (metric of Lobatchevski's non-euclidean geometry).
(a) Show that the Christoffel symbols of the Riemannian connection are: $\Gamma_{11}^{1}=\Gamma_{12}^{2}=\Gamma_{22}^{1}=0, \Gamma_{11}^{2}=\frac{1}{y}, \Gamma_{12}^{1}=\Gamma_{22}^{2}=-\frac{1}{y}$.
(b) Let $v_{0}=(0,1)$ be a tangent vector at point $(0,1)$ of $\mathbb{R}_{+}^{2}\left(v_{0}\right.$ is a unit vector on the $y$-axis with origin at $(0,1))$. Let $v(t)$ be the parallel transport of $v_{0}$ along the curve $x=t, y=1$. Show that $v(t)$ makes an angle $t$ with the direction of the $y$-axis, measured in the clockwise sense.
3. ([Car92] Ch. 3 \#5) Let $M$ be a Riemannian manifold and $X \in \mathcal{X}(M)$ (the set of all vector field of class $C^{\infty}$ on $\left.M\right)$. Let $p \in M$ and let $U \subset M$ be a neighborhood of $p$. Let $\varphi:(-\varepsilon, \varepsilon) \times U \rightarrow M$ be a differentiable mapping such that for any $q \in U$ the curve $t \rightarrow \varphi(t, q)$ is a trajectory of $X$ passing through $q$ at $t=0$ ( $U$ and $\varphi$ are given by the fundamental theorem for ordinary differential equations). $X$ is called a Killing field (or an infinitesimal isometry) if, for each $t_{0} \in(-\varepsilon, \varepsilon)$, the mapping $\varphi\left(t_{0}, \cdot\right)$ : $U \subset M \rightarrow M$ is an isometry. Prove that
(a) A vector field $v$ on $\mathbb{R}^{n}$ may be seen as a map $v: \mathbb{R}^{n} \rightarrow \mathbb{R}^{n}$; we say that the field is linear if $v$ is a linear map. A linear field on $\mathbb{R}^{n}$, defined by a matrix $A$, is a Killing field if and only if $A$ is anti-symmetric.
(b) Let $X$ be a Killing field on $M, p \in M$, and let $U$ be a normal neighborhood of $p$ on $M$. Assume that $p$ is a unique point of $U$ that satisfies $X(p)=0$. Then, in $U, X$ is tangent to the geodesic spheres centered at $p$.
(c) Let $X$ be a differentiable vector field on $M$ and let $f: M \rightarrow N$ be an isometry. Let $Y$ be a vector field on $N$ defined by $Y(f(p))=$ $d f_{p}(X(p)), p \in M$. Then $Y$ is a Killing field if and only if $X$ is also a Killing vector field.
(d) $X$ is Killing $\Leftrightarrow\left\langle\nabla_{Y} X, Z\right\rangle+\left\langle\nabla_{Z} X, Y\right\rangle=0$ for all $Y, Z \in \mathcal{X}(M)$ (the equation above is called the Killing equation).
(e) Let $X$ be a Killing field on $M$ with $X(q) \neq 0, q \in M$. Then there exists a system of coordinates $\left(x_{1}, \ldots, x_{n}\right)$ in a neighborhood of $q$, so that the coefficients $g_{i j}$ of the metric in this system coordinates do not depend on $x_{n}$.
4. ([Car92] Ch. 4 \#4) Let $M$ be a Riemannian manifold with the following property: given any two points $p, q \in M$, the parallel transport from $p$ to $q$ does not depend on the curve that joins $p$ to $q$. Prove that the curvature of $M$ is identically zero, that is, for all $X, Y, Z \in \mathcal{X}(M), R(X, Y) Z=0$.
5. Define covariant derivatives on differential forms.
6. ([Car92] Ch.4 \#8)(Schur's theorem) Let $M^{n}$ be a connected Riemannian manifold with $n \geq 3$. Suppose that $M$ is isotropic, that is, for each $p \in M$, the sectional curvature $K(p, \sigma)$ does not depend on $\sigma \subset T_{p} M$. Prove that $M$ has constant sectional curvature, that is, $K(p, \sigma)$ also does not depend on $p$.
7. ([Car92] Ch. 4 \#10) (Einstein manifolds) A Riemannian manifold $M^{n}$ is called an Einstein manifold if, for all $X, Y \in \mathcal{X}(M), \operatorname{Ric}(X, Y)=\lambda\langle X, Y\rangle$, where $\lambda: M \rightarrow \mathbb{R}$ is a real valued funciton. Prove that:
(a) If $M^{n}$ connected and Einstein, with $n \geq 3$, then $\lambda$ is constant on $M$.
(b) If $M^{3}$ is a connected Einstein manifold then $M^{3}$ has constant sectional curvature.
8. Show that the sectional curvature is the Gauss curvature.
9. ([Car92] Ch.4 \#6) Let $M$ be a Riemannian manifold. $M$ is a locally symmetric space if $\nabla R=0$, where $R$ is the curvature tensor of $M$.
(a) Let $M$ be a locally symmetric space and let $\gamma:[0, \ell) \rightarrow M$ be a geodesic of $M$. Let $X, Y, Z$ be parallel vector fields along $\gamma$. Prove that $R(X, Y) \mathrm{Z}$ is a parallel field along $\gamma$.
(b) Prove that if $M$ is locally symmetric, connected, and has dimension two, then $M$ has constant sectional curvature.
(c) Prove that if $M$ has constant (sectional) curvature, then $M$ is a locally symmetric space.
10. ([Car92] Ch.5 \#5) (Jacobi fields and conjugate points on locally symmetric spaces) Let $\gamma:[0, \infty) \rightarrow M$ be a geodesic in a locally symmetric space $M$ and let $v=\gamma^{\prime}(0)$ be its velocity at $p=\gamma(0)$. Define a linear transformation $K_{v}: T_{p} M \rightarrow T_{p} M$ by

$$
K_{v}(x)=R(v, x) v, x \in T_{p} M .
$$

(a) Prove that $K_{v}$ is self-adjoint.
(b) Choose an orthonormal basis $\left\{e_{1}, \ldots, e_{n}\right\}$ of $T_{p} M$ that diagonalizes $K_{v}$, that is,

$$
K_{v}\left(e_{i}\right)=\lambda_{i} e_{i}, \quad i=1, \ldots, n
$$

Extend the $e_{i}$ to fields along $\gamma$ by parallel transport. Show that, for all $t$,

$$
K_{\gamma^{\prime}(t)}\left(e_{i}(t)\right)=\lambda_{i} e_{i}(t)
$$

where $\lambda_{i}$ does not depend on $t$.
(c) Let $J(t)=\sum_{i} x_{i}(t) e_{i}(t)$ be a Jacobi field along $\gamma$. Show that the Jacobi equation is equivalent to the system

$$
\frac{d^{2} x_{i}}{d t^{2}}+\lambda_{i} x_{i}=0, \quad i=1, \ldots, n
$$

(d) Show that the conjugate points of $p$ along $\gamma$ are given by $\gamma\left(\pi k / \sqrt{\lambda_{i}}\right)$, where $k$ is a positive integer and $\lambda_{i}$ is a positive eigenvalue of $K_{v}$.
11. ([Car92] Ch.7 \#1) If $M, N$ are Riemannian manifolds such that the inclusion $i: M \hookrightarrow N$ is an isometric immersion, show by an example that the strict inequality of metrics $d_{M}>d_{N}$ can occur.
12. ([Car92] Ch.7 \#4) Consider the universal covering

$$
\pi: M \rightarrow \mathbb{R}^{2} \backslash\{(0,0)\}
$$

of the Euclidean plane minus the origin. Introduce the covering metric on $M$. (Note: For $\tilde{M}$ a covering space of $M$, it is possible to give the covering space a Riemannian structure such that the covering map $\pi: \tilde{M} \rightarrow M$ is a local isometry.) Show that $M$ is not complete and
not extendible, and that the Hopf-Rinow theorem is not true for $M$ (this shows that the definition of non-extendibility, though natural, is not a satisfactory one).
13. ([Car92] Ch. 7 \#5) A divergent curve in a Riemannian manifold $M$ is a differentiable mapping $\alpha:[0, \infty) \rightarrow M$ such that for any compact set $K \subset M$ there exists $t_{0} \in(0, \infty)$ with $\alpha(t) \notin K$ for all $t>t_{0}$ (that is, $\alpha$ "escapes" every compact set in $M$ ). Define the length of a divergent curve by

$$
\lim _{t \rightarrow \infty} \int_{0}^{t}\left|\alpha^{\prime}(t)\right| d t
$$

Prove that $M$ is complete if and only if the length of any divergent curve is unbounded.
14. ([Car92] Ch.7 \#6) A geodesic $\gamma:[0, \infty) \rightarrow M$ in a Riemannian manifold $M$ is called a ray starting from $\gamma(0)$ if it minimizes the distance between $\gamma(0)$ and $\gamma(s)$, for any $s \in(0, \infty)$. Assume that $M$ is complete, noncompact, and let $p \in M$. Show that $M$ contains a ray starting form $p$.
15. ([Car92] Ch. 7 \#12) A Riemannian manifold is said to be homogeneous if given $p, q \in M$ there exists an isometry of $M$ which takes $p$ into $q$. Prove that any homogeneous manifold is complete.
16. ([Car92] Ch. 8 \#1) Consider, on a neighborhood in $\mathbb{R}^{2}, n>2$ the metric

$$
g_{i j}=\frac{\delta_{i j}}{F^{2}}
$$

where $F \neq 0$ is a function of $\left(x_{1}, \ldots, x_{n}\right) \in \mathbb{R}^{n}$. Denote by $F_{i}=\frac{\partial F}{\partial x_{i}}$, $F_{i j}=\frac{\partial^{2} F}{\partial x_{i} \partial x_{j}}$, etc.
(a) Show that a necessary and sufficient condition for the metric to have constant curvature $K$ is

$$
\left\{\begin{array}{l}
F_{i j}=0, \quad i \neq j \\
F\left(F_{j j}+F_{i i}\right)=K+\sum_{i=1}^{n}\left(F_{i}\right)^{2}
\end{array}\right.
$$

(b) Use (a) to prove that the metric $g_{i j}$ has constant curvature $K$ if and only if

$$
F=G_{1}\left(x_{1}\right)+G_{2}\left(x_{2}\right)+\cdots+G_{n}\left(x_{n}\right),
$$

where

$$
G_{i}\left(x_{i}\right)=a x_{i}^{2}+b_{i} x_{i}+c_{i}
$$

and

$$
\sum_{i=1}^{n}\left(4 c_{i} a-b_{i}^{2}\right)=K
$$

(c) Put $a=K / 4, b_{i}=0, c_{i}=1 / n$ and obtain the formula of Riemann

$$
g_{i j}=\frac{\delta_{i j}}{\left(1+\frac{K}{4} \sum x_{i}^{2}\right)^{2}}
$$

for a metric $g_{i j}$ of constant curvature $K$. If $K<0$, the metric $g_{i j}$ is defined in a ball of radius of $\sqrt{\frac{4}{-K}}$.
(d) If $K>0$, the metric is defined on all of $\mathbb{R}^{n}$. Show that such a metric on $\mathbb{R}^{n}$ is not complete.
17. ([Car92] Ch.6 \#1) Let $M_{1}$ and $M_{2}$ be Riemannian manifolds, and consider the product $M_{1} \times M_{2}$, with the product metric. Let $\nabla^{1}$ be the Riemannian connection of $M_{1}$ and let $\nabla^{2}$ be the Riemannian connection of $M_{2}$.
(a) Show that the Riemannian connection $\nabla$ of $M_{1} \times M_{2}$ is given by $\nabla_{Y_{1}+Y_{2}}\left(X_{1}+X_{2}\right)=\nabla_{Y_{1}}^{1} X_{1}+\nabla_{Y_{2}}^{2} X_{2}$, for $X_{1}, Y_{1} \in \mathcal{X}\left(M_{1}\right), X_{2}, Y_{2} \in$ $\mathcal{X}\left(M_{2}\right)$.
(b) For every $p \in M_{1}$, the set $\left(M_{2}\right)_{p}=\left\{(p, q) \in M_{1} \times M_{2} ; q \in M_{2}\right\}$ is a submanifold of $M_{1} \times M_{2}$, naturally diffeomorphic to $M_{2}$. Prove that $\left(M_{2}\right)_{p}$ is a totally geodesic submanifold of $M_{1} \times M_{2}$.
(c) Let $\sigma(x, y) \subset T_{(p, q)}\left(M_{1} \times M_{2}\right)$ be a plane such that $x \in T_{p} M_{1}$ and $y \in T_{q} M_{2}$. Show that $K(\sigma)=0$.
18. ([Car92] Ch.6 \#2) Show that $\mathbf{x}: \mathbb{R}^{2} \rightarrow \mathbb{R}^{4}$ given by

$$
\mathbf{x}(\theta, \varphi)=\frac{1}{\sqrt{2}}(\cos \theta, \sin \theta, \cos \varphi, \sin \varphi),(\theta, \varphi) \in \mathbb{R}^{2}
$$

is an immersion of $\mathbb{R}^{2}$ into the unit sphere $S^{3}(1) \subset \mathbb{R}^{4}$, whose image $\mathbf{x}\left(\mathbb{R}^{2}\right)$ is a torus $\mathbb{T}^{2}$ with sectional curvature zero in the induced metric.
19. ([Car92] Ch.6 \#4) Let $N_{1} \subset M_{1}, N_{2} \subset M_{2}$ be totally geodesic submanifolds of the Riemannian manifolds $M_{1}$ and $M_{2}$ respectively. Prove that $N_{1} \times N_{2}$ is a totally geodesic submanifold of the product $M_{1} \times M_{2}$ with the product metric.
20. ([Car92] Ch. 6 \#5) Prove that the sectional curvature of the Riemannian manifold $S^{2} \times S^{2}$ with the product metric, where $S^{2}$ is the unit sphere in $\mathbb{R}^{3}$, is non-negative. Find a totally geodesic, flat torus, $\mathbb{T}^{2}$, embedded in $S^{2} \times S^{2}$.
21. ([Car92] Ch. 6 \#8) (The Clifford torus) Consider the immersion $\mathbf{x}: \mathbb{R}^{2} \rightarrow$ $\mathbb{R}^{4}$ given in the above exercise 18.
(a) Show that the vectors

$$
e_{1}=(-\sin \theta, \cos \theta, 0,0), e_{2}=(0,0,-\sin \varphi, \cos \varphi)
$$

form an orthonormal basis of the tangent space, and that the vectors

$$
\begin{aligned}
& n_{1}=\frac{1}{\sqrt{2}}(\cos \theta, \sin \theta, \cos \varphi, \sin \varphi) \\
& n_{2}=\frac{1}{\sqrt{2}}(-\cos \theta,-\sin \theta, \cos \varphi, \sin \varphi)
\end{aligned}
$$

form an orthonormal basis of the normal space.
(b) Use the fact that

$$
\left\langle S_{n_{k}}\left(e_{i}\right), e_{j}\right\rangle=-\left\langle\bar{\nabla}_{e_{i}} n_{k}, e_{j}\right\rangle=\left\langle\bar{\nabla}_{e_{i}} e_{j}, n_{k}\right\rangle,
$$

where $\bar{\nabla}$ is the covariant derivative (that is, the usual derivative) of $\mathbb{R}^{4}$, and $i, j, k=1,2$, to establish that the matrices of $S_{n_{1}}$ and $S_{n_{2}}$ with respect to the basis $\left\{e_{1}, e_{2}\right\}$ are

$$
S_{n_{1}}=\left[\begin{array}{cc}
-1 & 0 \\
0 & -1
\end{array}\right], \quad S_{n_{2}}=\left[\begin{array}{cc}
1 & 0 \\
0 & -1
\end{array}\right]
$$

(c) From the above exercise $18, \mathbf{x}$ is an immersion of the torus $\mathbb{T}^{2}$ into $S^{3}(1)$ (the Clifford torus). Show that $\mathbf{x}$ is a minimal immersion.
22. ([Car92] Ch.6 \#11) Let $f: \bar{M}^{n+1} \rightarrow \mathbb{R}$ be a differentiable function. Define the Hessian, Hess $f$ of $f$ at $p \in \bar{M}$ as the linear operator

$$
\text { Hess } f: T_{p} \bar{M} \rightarrow T_{p} \bar{M},(\operatorname{Hess} f) Y=\bar{\nabla}_{Y} \operatorname{grad} f, Y \in T_{p} \bar{M}
$$

where $\bar{\nabla}$ is the Riemannian connection of $\bar{M}$. Let $a$ be a regular value of $f$ and let $M^{n} \subset \bar{M}^{n+1}$ be the hypersuperface in $\bar{M}$ defined by $M=\{p \in$ $\bar{M} \mid f(p)=a\}$. Prove that:
(a) The Laplacian $\bar{\Delta} f$ is given by

$$
\bar{\Delta} f=\operatorname{trace}(\operatorname{Hess} f)
$$

(b) If $X, Y \in \mathcal{X}(M)$, then

$$
\langle(\operatorname{Hess} f) Y, X\rangle=\langle Y,(\operatorname{Hess} f) X\rangle
$$

Conclude that Hess $f$ is self-adjoint, hence determines a symmetric bilinear form on $T_{p} \bar{M}, p \in \bar{M}$, given by $($ Hess $f)(X, Y)=\langle($ Hess $f) X, Y\rangle$, $X, Y \in T_{p} \bar{M}$.
(c) The mean curvature $H$ of $M \subset \bar{M}$ is given by

$$
n H=-\operatorname{div}\left(\frac{\operatorname{grad} f}{|\operatorname{grad} f|}\right) .
$$

(d) Observe that every embedded hypersurface $M^{n} \subset \bar{M}^{n+1}$ is locally the inverse image of a regular value. Conclude from (c) that the mean curvature $H$ of such a hypersuperface is given by

$$
H=-\frac{1}{n} \operatorname{div} H,
$$

where $N$ is an appropriate local extension of the unit normal vector field on $M^{n} \subset \bar{M}^{n+1}$.
23. ([Car92] Ch. 8 \#8) (Riemannian submersions) A differentiable mapping $f: \bar{M}^{n+k} \rightarrow M^{n}$ is called a submersion if $f$ is surjective, and for all $\bar{p} \in \bar{M}$, $d f_{\bar{p}}: T_{\bar{p}} \bar{M} \rightarrow T_{f(\bar{p})} M$ has rank $n$. In this case, for all $p \in M$, the fiber $f^{-1}(p)=F_{p}$ is a submanifold of $\bar{M}$ and a tangent vector of $\bar{M}$, tangent to some $F_{p}, p \in M$, is called a vertical vector of the submersion. If, in addition, $\bar{M}$ and $M$ have Riemannian metrics, the submersion $f$ is said to be Riemannian if, for all $p \in \bar{M}, d f_{p}: T_{p} \bar{M} \rightarrow T_{f(p)} M$ preserves lengths of vectors orthogonal to $F_{p}$. Show that:
(a) If $M_{1} \times M_{2}$ is the Riemannian product, then the natural projections $\pi_{i}: M_{1} \times M_{2} \rightarrow M_{i}, i=1,2$ are Riemannian submersions.
(b) Let the tangent bundle $T M$ be given the Riemannian metric as:

$$
\langle V, W\rangle_{(p, v)}=\langle d \pi(V), d \pi(W)\rangle_{p}+\left\langle\frac{D v}{d t}(0), \frac{D w}{d s}(0)\right\rangle_{p}
$$

for $(p, v) \in T M, V, W$ tangent vectors at $(p, v)$ in $T M$ where $V=$ $\alpha^{\prime}(0), W=\beta^{\prime}(0)$ for curves $\alpha, \beta$ chosen such that $\alpha(t)=(p(t), v(t))$, $\beta(t)=(q(s), w(s)), p(0)=q(0)=0, v(0)=w(0)=v(c f .[C a r 92]$ Ch. 3 \#2). Show that the projection $\pi: T M \rightarrow M$ is a Riemannian submersion.
24. ([Car92] Ch.8 \#9) (Conneciton of a Riemannian submersion) Let $f: \bar{M} \rightarrow$ $M$ be a Riemannian submersion. A vector $\bar{x} \in T_{\bar{p}} \bar{M}$ is horizontal if it is orthogonal to the fiber. The tangent space $T_{\bar{p}} \bar{M}$ then admits a decomposition $T_{\bar{p}} \bar{M}=\left(T_{\bar{p}} \bar{M}\right)^{h} \oplus\left(T_{\bar{p}} \bar{M}\right)^{v}$, where $\left(T_{\bar{p}} \bar{M}\right)^{h}$ and $\left(T_{\bar{p}} \bar{M}\right)^{v}$ denote the subspaces of horizontal and vertical vectors, respectively. If $X \in \mathcal{X}(M)$, the horizontal lift $\bar{X}$ of $X$ is the horizontal field defined by $d f_{\bar{p}}(\bar{X}(\bar{p}))=X(f(p))$.
(a) Show that $\bar{X}$ is differentiable.
(b) Let $\nabla$ and $\bar{\nabla}$ be the Riemannian connections of $M$ and $\bar{M}$ respectively. Show that

$$
\bar{\nabla}_{\bar{X}} \bar{Y}=\overline{\nabla_{X} Y}+\frac{1}{2}[\bar{X}, \bar{Y}]^{v}, \quad X, Y \in \mathcal{X}(M)
$$

where $Z^{v}$ is the vertical component of $Z$.
(c) $[\bar{X}, \bar{Y}]^{v}(\bar{p})$ depends only on $\bar{X}(\bar{p})$ and $\bar{Y}(\bar{p})$.
25. ([Car92] Ch.8\#10) (Curvature of a Riemannian submersion) Let $f: \bar{M} \rightarrow$ $M$ be a Riemannian submersion. Let $X, Y, Z, W \in \mathcal{X}(M), \bar{X}, \bar{Y}, \bar{Z}, \bar{W}$ be their horizontal lifts, and let $R$ and $\bar{R}$ be the curvature tensors of $M$ and $\bar{M}$ respectively. Prove that:
(a)

$$
\begin{aligned}
\langle\bar{R}(\bar{X}, \bar{Y}) \bar{Z}, \bar{W}\rangle & =\langle R(X, Y) Z, W\rangle-\frac{1}{4}\left\langle[\bar{X}, \bar{Z}]^{v},[\bar{Y}, \bar{W}]^{v}\right\rangle \\
& +\frac{1}{4}\left\langle[\bar{Y}, \bar{Z}]^{v},[\bar{X}, \bar{W}]^{v}\right\rangle-\frac{1}{2}\left\langle[\bar{Z}, \bar{W}]^{v},[\bar{X}, \bar{Y}]^{v}\right\rangle
\end{aligned}
$$

(b) $K(\sigma)=\bar{K}(\bar{\sigma})+\frac{3}{4}\left|[\bar{X}, \bar{Y}]^{v}\right|^{2} \geq \bar{K}(\bar{\sigma})$, where $\sigma$ is the plane generated by the orthonormal vectors $X, Y \in \mathcal{X}(M)$ and $\bar{\sigma}$ is the plane generated by $\bar{X}, \bar{Y}$.
26. ([Car92] Ch.8 \#11) (The complex projective space) Let

$$
\mathbb{C}^{n+1} \backslash\{0\}=\left\{\left(z_{0}, \ldots, z_{n}\right)=\mathrm{Z} \neq 0 \mid z_{j}=x_{j}+i y_{j}, j=0, \ldots, n\right\}
$$

be the set of all non-zero $(n+1)$-tuples of complex numbers $z_{j}$. Define equivalence relation on $\mathbb{C}^{n+1} \backslash\{0\}:\left(z_{0}, \ldots, z_{n}\right) \sim W=\left(w_{0}, \ldots, w_{n}\right)$ if $z_{j}=\lambda w_{j}, \lambda \in \mathbb{C}, \lambda \neq 0$. The equivalence class of $Z$ will be denoted by [ $Z$ ] (the complex line passing through the origin and through $Z$ ). The set of such classes is called, by analogy with the real case, the complex projective space $\mathbb{P}^{n}(\mathbb{C})$ of complex dimension $n$.
(a) Show that $\mathbb{P}^{n}(\mathbb{C})$ has a differentiable structure of a manifold of real dimension $2 n$ and that $\mathbb{P}^{1}(\mathbb{C})$ is diffeomorphic to $\mathrm{S}^{2}$.
(b) Let $(Z, W)=z_{0} \overline{w_{0}}+\cdots+z_{n} \overline{w_{n}}$ be the hermitian product on $\mathbb{C}^{n+1}$, where the bar denotes complex conjugation. Identify $\mathbb{C}^{n+1} \approx \mathbb{R}^{2 n+2}$ by putting $z_{j}=x_{j}+i y_{j}=\left(x_{j}, y_{j}\right)$. Show that

$$
\mathrm{S}^{2 n+1}=\left\{N \in \mathbb{C}^{n+1} \approx \mathbb{R}^{2 n+2} \mid(N, N)=1\right\}
$$

is the unit sphere in $\mathbb{R}^{2 n+2}$.
(c) Show that the equivalence relation $\sim$ induces on $\mathrm{S}^{2 n+1}$ the following equivalence relation: $\mathrm{Z} \sim W$ if $e^{i \theta} Z=W$. Establish that there exists a differentiable map (the Hopf fibering) $f: S^{2 n+1} \rightarrow \mathbb{P}^{n}(\mathbb{C})$
such that

$$
\begin{aligned}
& f^{-1}([Z]) \\
& \quad=\left\{e^{i \theta} N \in \mathrm{~S}^{2 n+1} \mid N \in[Z] \cap \mathrm{S}^{2 n+1}, 0 \leq \theta \leq 2 \pi\right\} \\
& \quad=[Z] \cap \mathrm{S}^{2 n+1} .
\end{aligned}
$$

(d) Show that $f$ is a submersion.
27. ([Car92] Ch. 8 \#12) (Curvature of the complex projective space) Define a Riemannian metric on $\mathbb{C}^{n+1} \backslash\{0\}$ in the following way: If $Z \in \mathbb{C}^{n+1} \backslash\{0\}$ and $V, W \in T_{Z}\left(\mathbb{C}^{n+1} \backslash\{0\}\right)$,

$$
\langle V, W\rangle_{Z}=\frac{\operatorname{Real}(V, W)}{(Z, Z)}
$$

Observe that the metric $\langle$,$\rangle restricted to \mathbb{S}^{2 n+1} \subset \mathbb{C}^{n+1} \backslash\{0\}$ coincides with the metric induced from $\mathbb{R}^{2 n+2}$.
(a) Show that, for all $0 \leq \theta \leq 2 \pi, e^{i \theta}: \mathrm{S}^{2 n+1} \rightarrow \mathrm{~S}^{2 n+1}$ is an isometry, and that, therefore, it is possible to define a Riemannian metric on $\mathbb{P}^{n}(\mathbb{C})$ in such a way that the submersion $f$ is Riemannian.
(b) Show that, in this metric, the sectional curvature of $\mathbb{P}^{n}(\mathbb{C})$ is given by

$$
K(\sigma)=1+3 \cos ^{2} \varphi,
$$

where $\sigma$ is generated by the orthonormal pair $X, Y, \cos \varphi=\langle\bar{X}, i \bar{Y}\rangle$, and $\bar{X}, \bar{Y}$ are the horizontal lifts of $X$ and $Y$, respectively. In particular, $1 \leq K(\sigma) \leq 4$.


[^0]:    ${ }^{1}$ A major philosophical question in Riemannian geometry, essentially posed by Riemann, is the isometric imbedding problem: for any $(M, g)$, does there exists a $C^{\infty}$ imbedding $\iota: M \rightarrow \mathbb{R}^{N}$ for some $N \in \mathbb{N}$ such that

    $$
    g=\iota^{*} g_{\mathbb{R}^{N}} ?
    $$

    The question was finally answered positively by John Nash around 1954-1956 (1954: $C^{1}$ isometric imbeddings; 1956: $C^{\infty}$ isometric embeddings). Nash's proof opened a new era of non-linear analysis and in particular lead to his celebrated inverse function theorem on Fréchet spaces. We will not discuss it in this beginning stage. Will come back to some aspects on it in the later part of this book.

[^1]:    ${ }^{2}$ Different extensions $X$ lead to different integral curves and flows $\phi_{t}$. Even if we take a non-trivial scaling $h X$ of one extension $X$ with $h(p)=1$ so that we get the same integral curves, they must have different parametrizations and the flows are different. Thus in general we get different values of $L_{X} Y$ at $p$.

[^2]:    ${ }^{3}$ In Einstein's general theory of relativity in 1915, curvatures of spacetime are related by Einstein's field equation:

    $$
    R_{i j}-\frac{s}{2} g_{i j}=T_{i j},
    $$

